





Search for Charged Lepton Flavor Violation with Muons

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Outline



- Muon LFV for BSM search
- Muon LFV processes
 - $\mu \rightarrow e \gamma$
 - $\mu \rightarrow e e e$
 - $\mu \rightarrow e$ conversion
- Other muon LFV processes
- Muon sources in the world
- COMET@J-PARC
 - COMET Phase-II
 - COMET Phase-I
- Timeline
- Summary

COMET official Logo



Muon LFV for BSM Search

EFT with LFV



$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)}$$

dimension 6

Λ is the energy scale of new physics
 C_{NP} is the coupling constant.

EFT with LFV



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dimension 6

ex : $\mu \rightarrow e \gamma$ $B < 4.2 \times 10^{-13}$ from MEG(2016)

$$\frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)} \rightarrow \frac{C_{\mu e}}{\Lambda^2} \bar{e}_L \sigma^{\rho\nu} \mu_R \Phi F_{\rho\nu}$$

$$\Lambda > 2 \times 10^5 \text{ TeV} \times (C_{\mu e})^{\frac{1}{2}} .$$

Λ is the energy scale of new physics
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$$\Lambda > O(10^3) \text{ TeV with } C_{\mu e} \sim O(1)$$

or

$$C_{\mu e} \sim O(10^{-9}) \text{ with } \Lambda < O(1) \text{ TeV}$$

EFT with LFV



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$$\Lambda > O(10^3) \text{ TeV with } C_{\mu e} \sim O(1)$$

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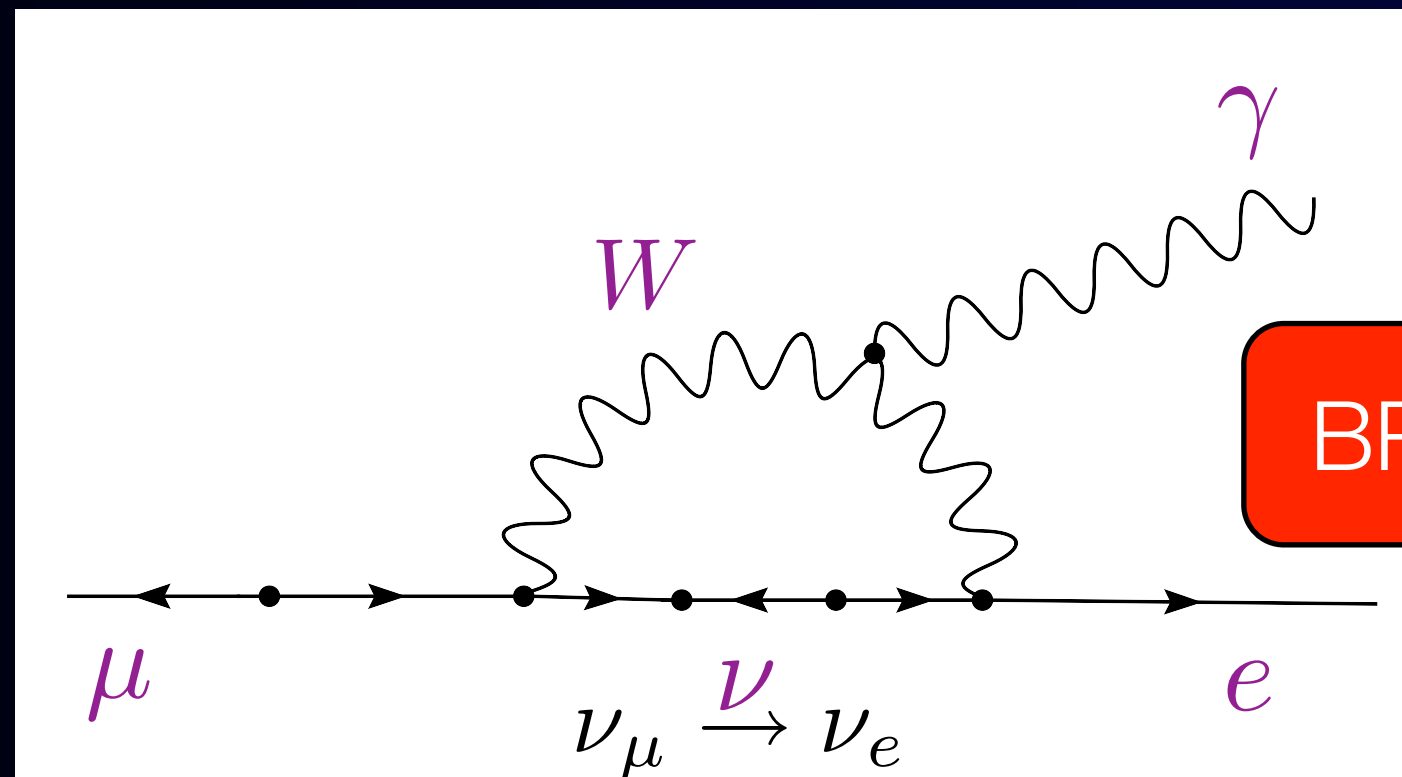
$$C_{\mu e} \sim O(10^{-9}) \text{ with } \Lambda < O(1) \text{ TeV}$$

can explore very high energy scale that
accelerators cannot directly reach.

SM Contribution to CLFV



SM
Contribution
with neutrino
mixing



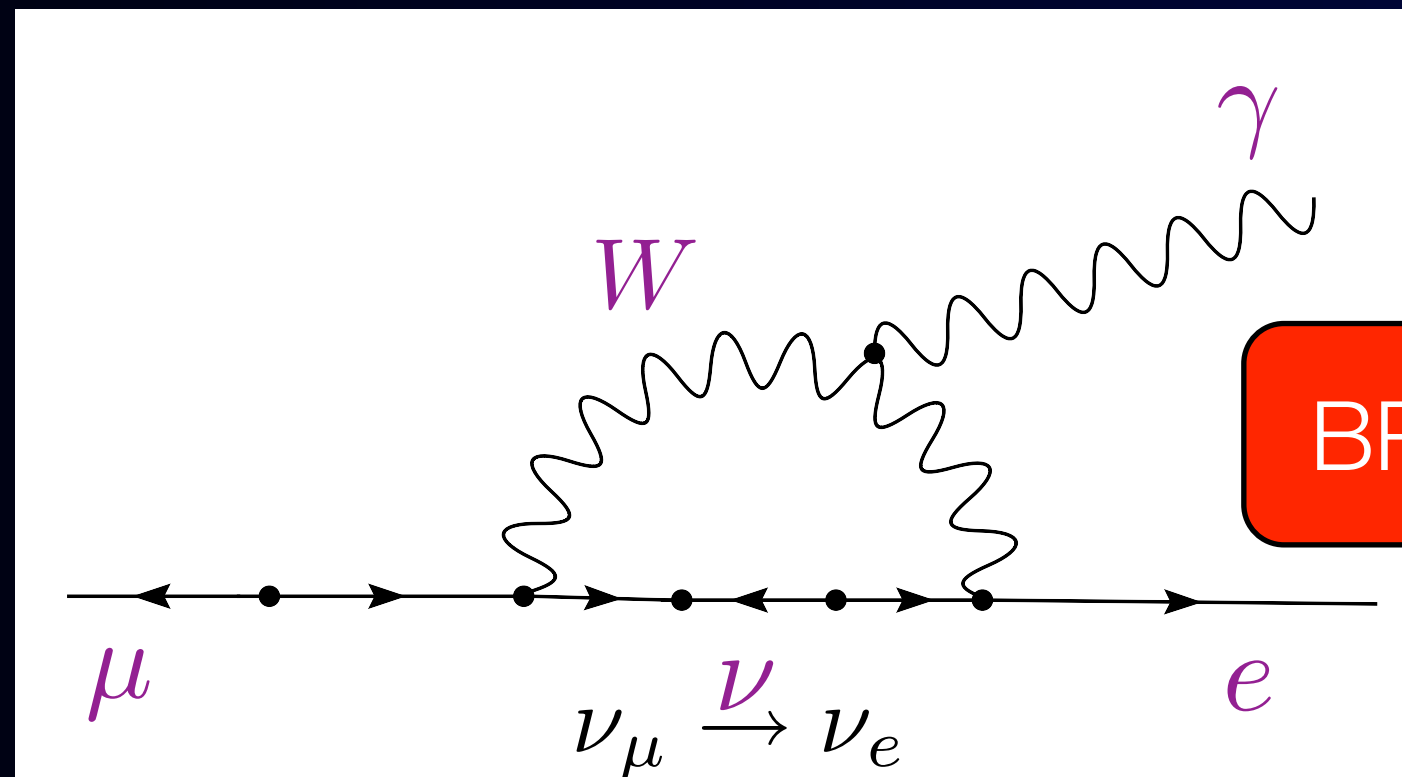
BR $\sim O(10^{-54})$

$$B(\mu \rightarrow e \gamma) = \frac{3\alpha}{32\pi} \left| \sum_l (V_{MNS})_{\mu l}^* (V_{MNS})_{el} \frac{m_{\nu_l}^2}{M_W^2} \right|^2$$

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BR ~ O(10⁻⁵⁴)

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Observation of CLFV would indicate a clear signal of physics beyond the SM with massive neutrinos.

Model Dependent CLFV



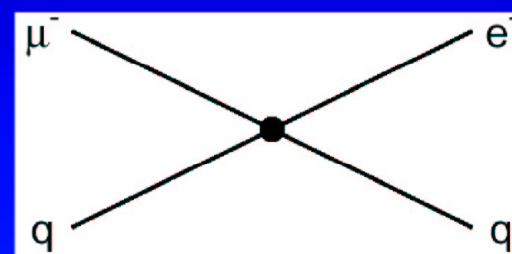
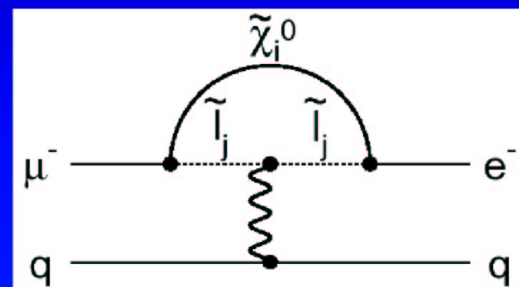
Model Dependent CLFV



Sensitivity to Different Muon Conversion Mechanisms



Supersymmetry
Predictions at 10^{-15}



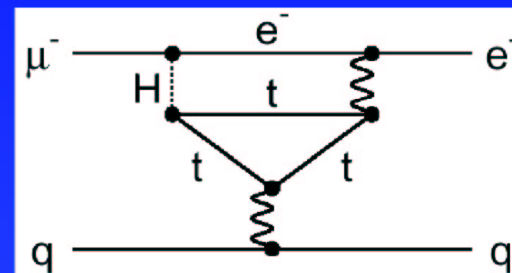
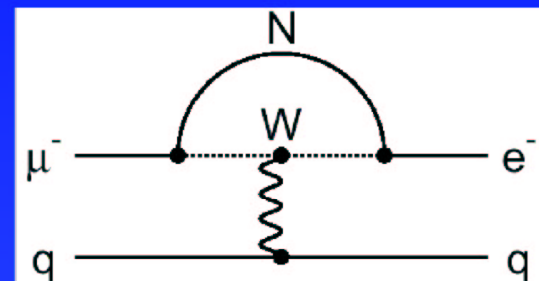
Compositeness

$$\Lambda_c = 3000 \text{ TeV}$$

Heavy Neutrinos

$$|U_{\mu N}^* U_{eN}|^2 =$$

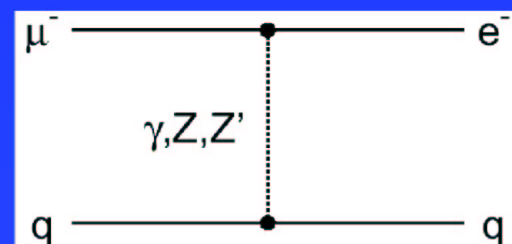
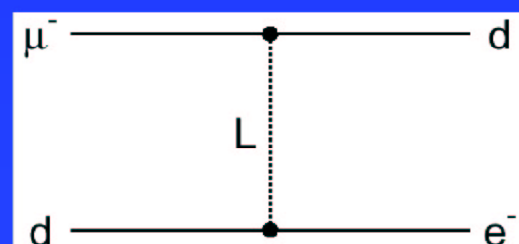
$$8 \times 10^{-13}$$



Second Higgs
doublet

$$g_{H\mu e} = 10^{-4} \times g_{H\mu\mu}$$

Leptoquarks



Heavy Z' ,
Anomalous Z
coupling

$$M_{Z'} = 3000 \text{ TeV}/c^2$$

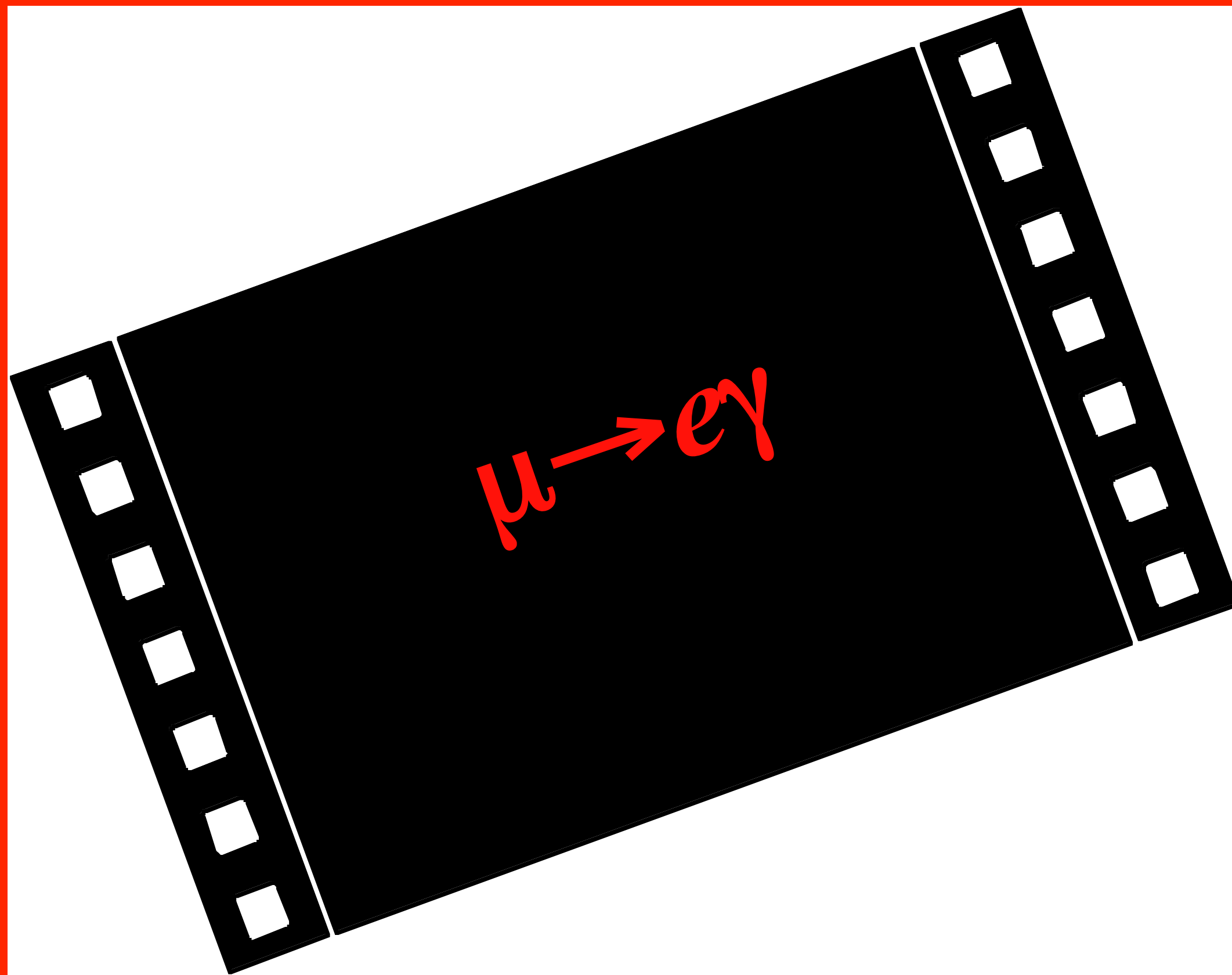
$$B(Z \rightarrow \mu e) < 10^{-17}$$

$$M_L =$$

$$3000 (\lambda_{\mu d} \lambda_{e d})^{1/2} \text{ TeV}/c^2$$

After W. Marciano

Muon CLFV Processes



$$\mu \rightarrow e\gamma$$



$$\mu \rightarrow e \gamma$$



- **Event Signature**
 - $E_e = E_\gamma = m_\mu/2$ (=52.8 MeV)
 - angle $\theta_{\mu e}=180$ (back-to-back)
 - time coincidence
- **Backgrounds**
 - prompt physics backgrounds
 - radiative muon decay
 $\mu \rightarrow e \nu \nu \gamma$
 - accidental backgrounds

$\mu \rightarrow e \gamma$



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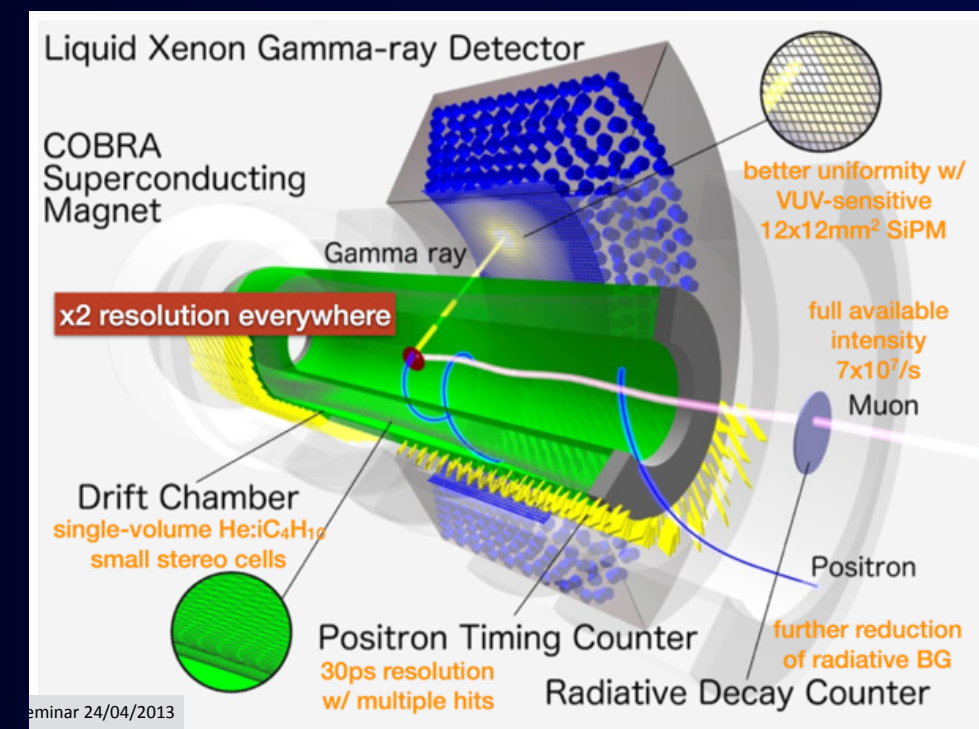
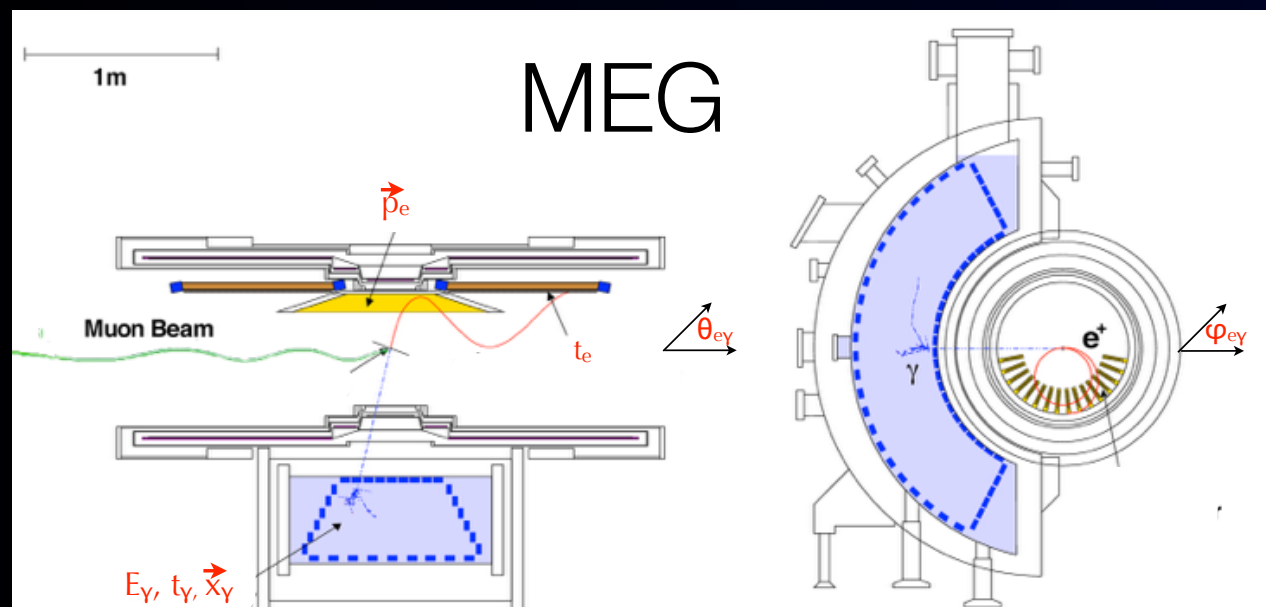
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Final MEG result (2016)

$$B(\mu^+ \rightarrow e^+ \gamma) < 4.2 \times 10^{-13}$$

MEG II

goal $\sim 4 \times 10^{-14}$
2018-2020



$\mu \rightarrow eee$

$\mu \rightarrow eee$



$\mu \rightarrow eee$



- **Event Signature**

- $\Sigma E_e = m_\mu$
- $\Sigma \mathbf{P}_e = 0$ (vector sum)
- common vertex
- time coincidence

- **Backgrounds**

- physics backgrounds
 - $\mu \rightarrow e \nu \nu e$ decay
- accidental backgrounds

$\mu \rightarrow eee$



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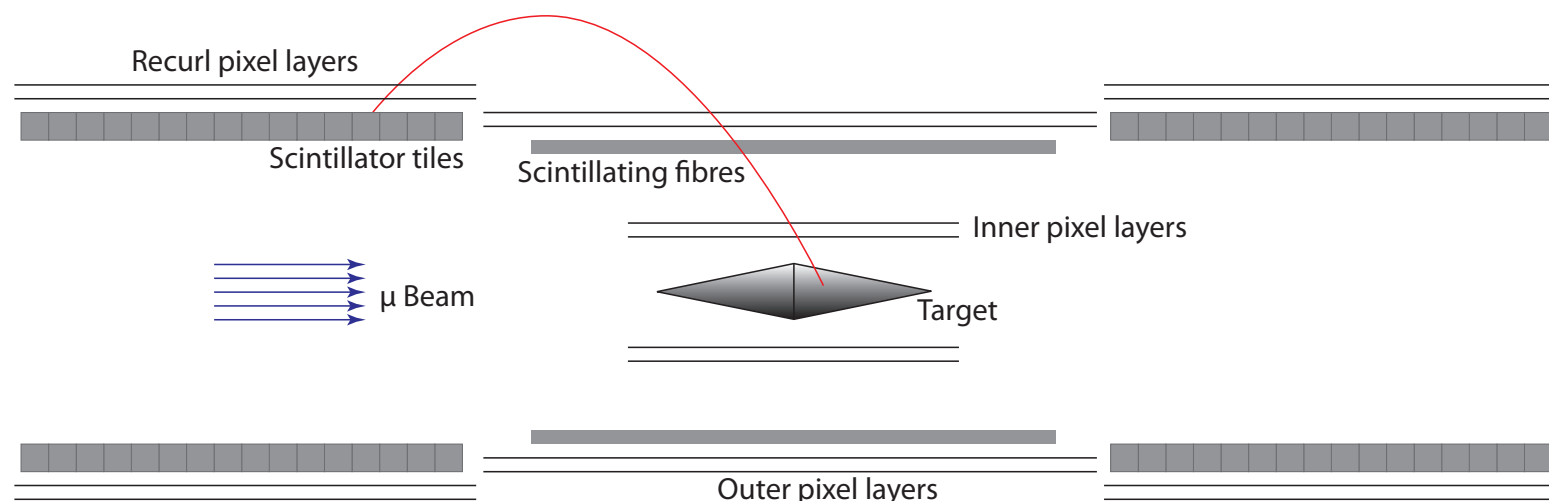
- **Backgrounds**

- physics backgrounds
 - $\mu \rightarrow e\nu\bar{\nu}e$ decay
- accidental backgrounds

Mu3e (@PSI)

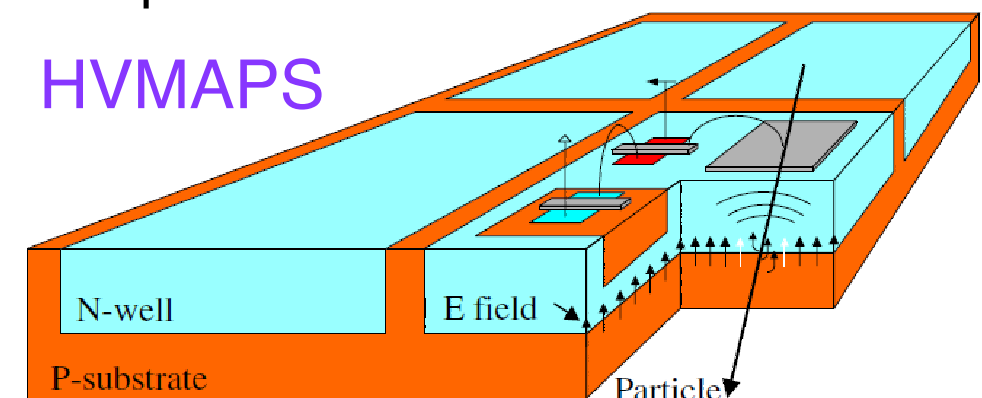
- current limit : 1×10^{-12}
- Stage-I (2020 -)
 - $B \sim 10^{-15}$ at $\pi E5$
- Stage-2
 - $B < 10^{-16}$ at new muon source

Mu3e



50μm thickness

HVMAPS





$\mu \rightarrow e$ conversion
in
a muonic atom

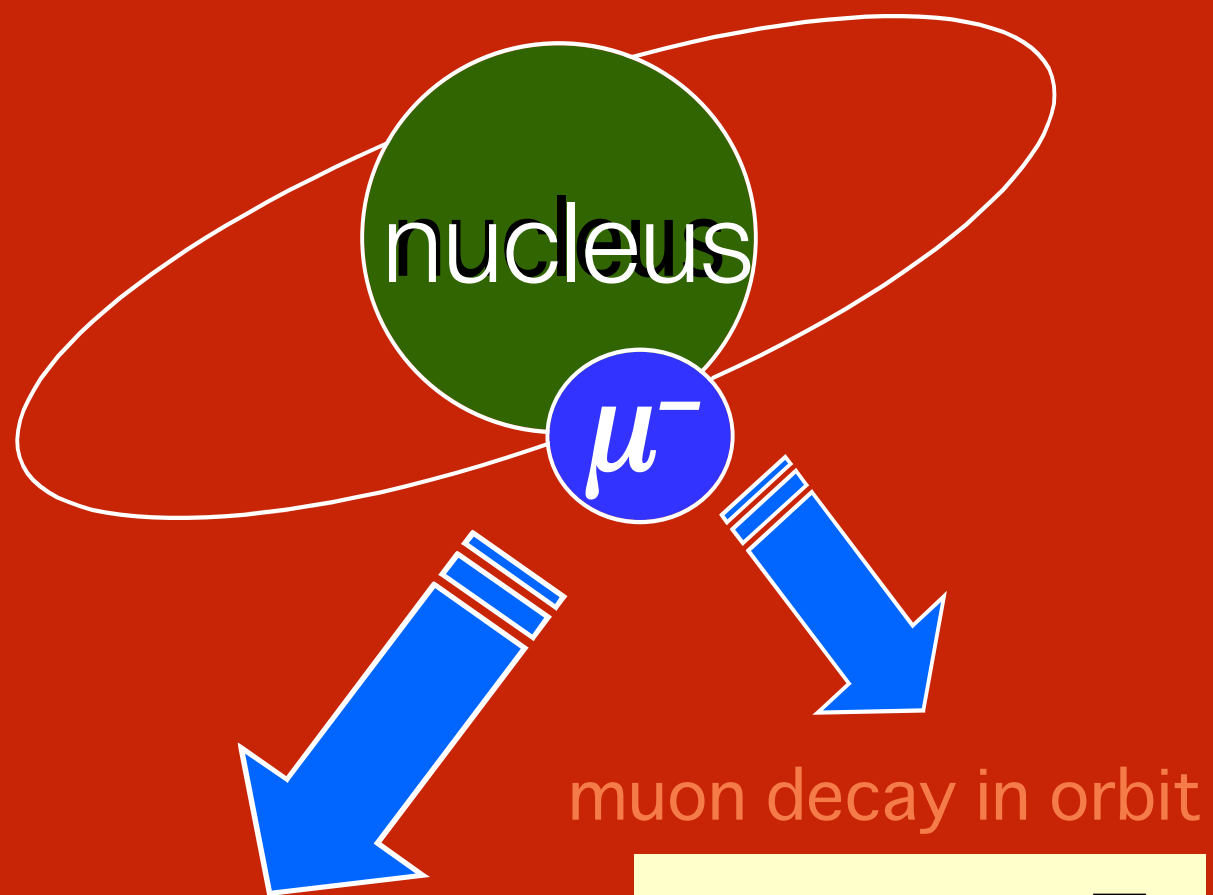
What is $\mu \rightarrow e$ Conversion ?



What is $\mu \rightarrow e$ Conversion ?



1s state in a muonic atom



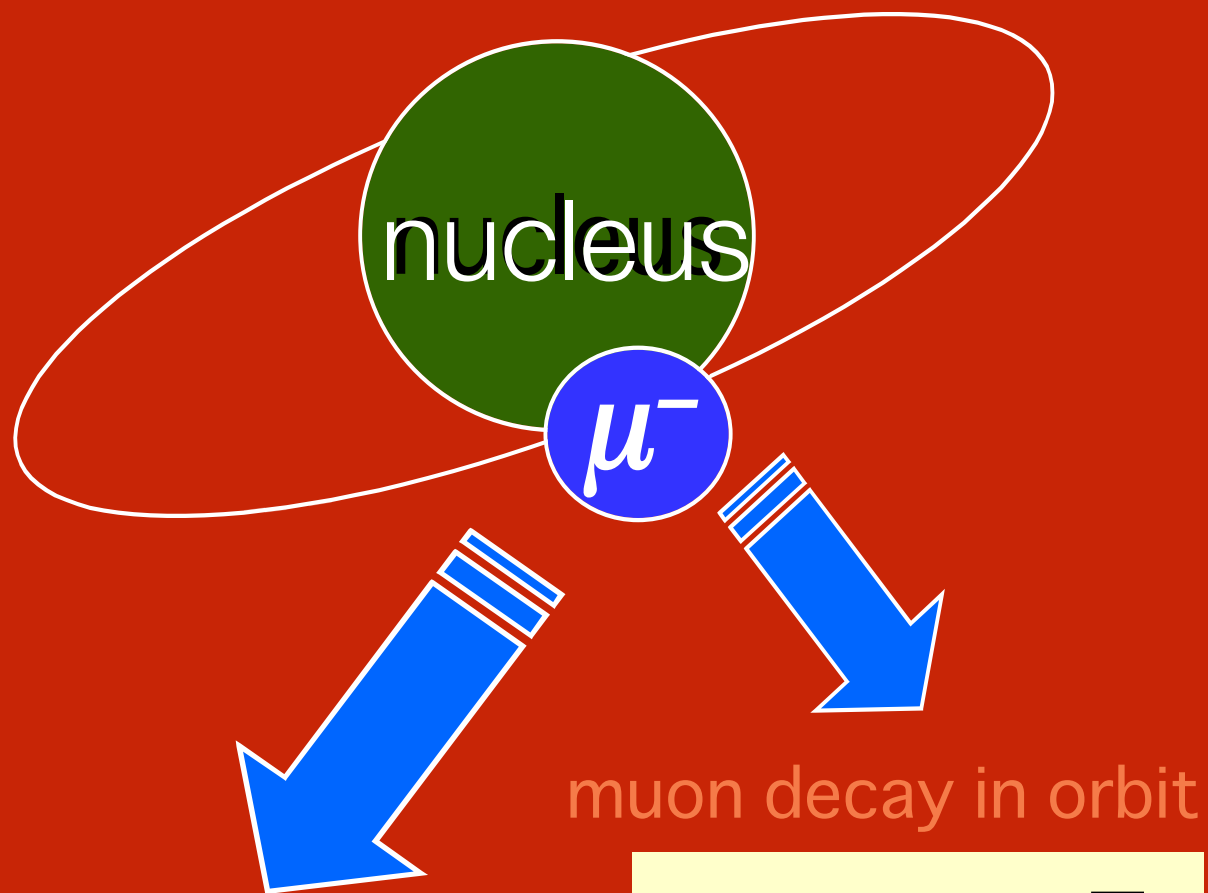
$$\mu^- \rightarrow e^- \nu \bar{\nu}$$

nuclear muon capture

$$\mu^- + (A, Z) \rightarrow \nu_\mu + (A, Z - 1)$$

What is $\mu \rightarrow e$ Conversion ?

1s state in a muonic atom



muon decay in orbit

$$\mu^- \rightarrow e^- \nu \bar{\nu}$$

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Neutrino-less muon
nuclear capture

$$\mu^- + (A, Z) \rightarrow e^- + (A, Z)$$

Event Signature :

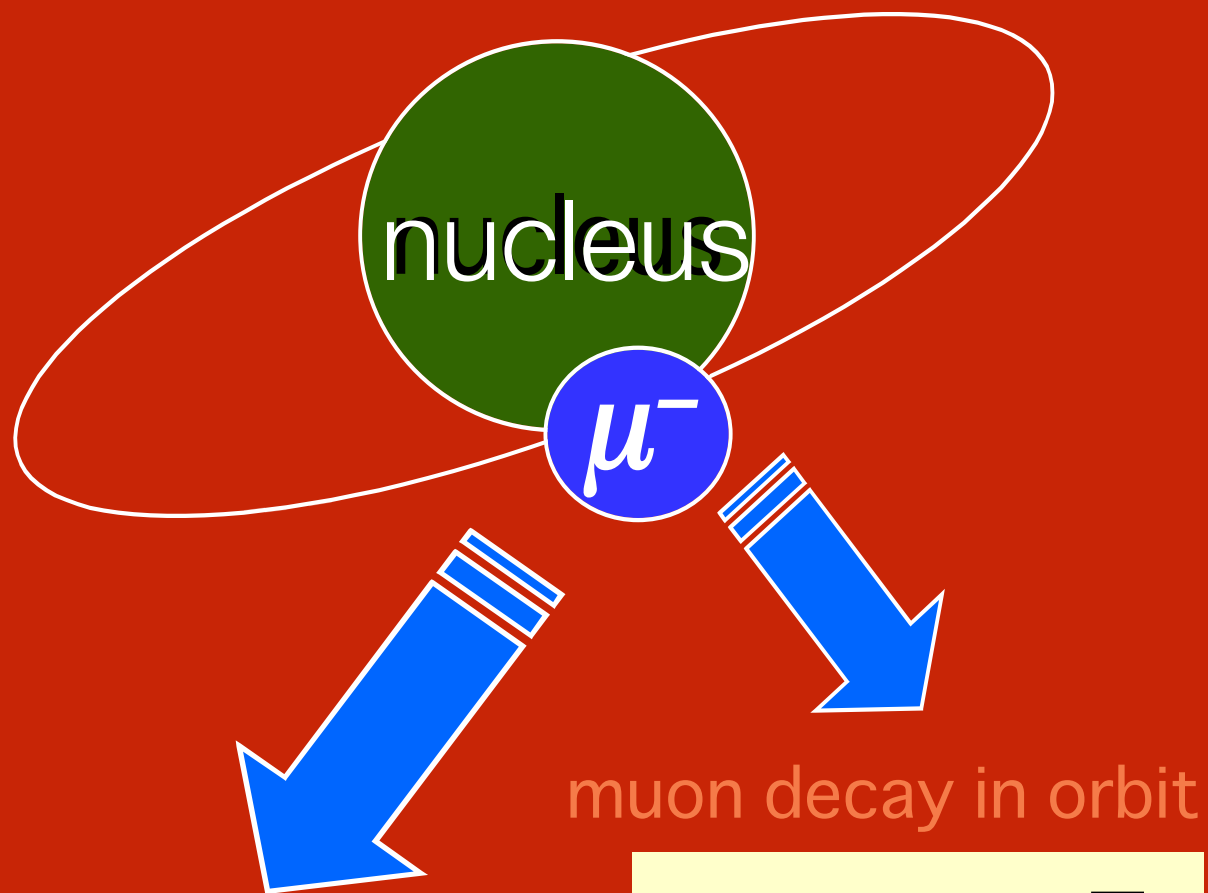
a single mono-energetic
electron of 105 MeV

Backgrounds:

- (1) physics backgrounds
- (2) beam-related backgrounds
- (3) cosmic rays, false tracking

What is $\mu \rightarrow e$ Conversion ?

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COMET explained later

Experimental Comparison



	Beam	background	challenge	beam intensity
$\mu \rightarrow e\gamma$	continuous beam	accidentals	detector resolution	limited
$\mu \rightarrow eee$	continuos beam	accidentals	detector resolution	limited
μ -e conversion	pulsed beam	beam-related	beam background	no limitation

if found,

Effective Field Theory for $\mu \rightarrow e$ Conversion



Effective Field Theory for $\mu \rightarrow e$ Conversion



$$\begin{aligned}
 \mathcal{L}_{\mu A \rightarrow e A}(\Lambda_{expt}) = & -\frac{4G_F}{\sqrt{2}} \sum_{N=p,n} \left[m_\mu (C_{DL} \bar{e}_R \sigma^{\alpha\beta} \mu_L F_{\alpha\beta} + C_{DR} \bar{e}_L \sigma^{\alpha\beta} \mu_R F_{\alpha\beta}) \right. \\
 & \text{scalar} \quad + \left(\tilde{C}_{SL}^{(NN)} \bar{e} P_L \mu + \tilde{C}_{SR}^{(NN)} \bar{e} P_R \mu \right) \bar{N} N \\
 & \text{pseudo-scalar} \quad + \left(\tilde{C}_{P,L}^{(NN)} \bar{e} P_L \mu + \tilde{C}_{P,R}^{(NN)} \bar{e} P_R \mu \right) \bar{N} \gamma_5 N \\
 & \text{vector} \quad + \left(\tilde{C}_{VL}^{(NN)} \bar{e} \gamma^\alpha P_L \mu + \tilde{C}_{VR}^{(NN)} \bar{e} \gamma^\alpha P_R \mu \right) \bar{N} \gamma_\alpha N \\
 & \text{axial-vector} \quad + \left(\tilde{C}_{A,L}^{(NN)} \bar{e} \gamma^\alpha P_L \mu + \tilde{C}_{A,R}^{(NN)} \bar{e} \gamma^\alpha P_R \mu \right) \bar{N} \gamma_\alpha \gamma_5 N \\
 & \text{tensor} \quad + \left. \left(\tilde{C}_{T,L}^{(NN)} \bar{e} \sigma^{\alpha\beta} P_L \mu + \tilde{C}_{T,R}^{(NN)} \bar{e} \sigma^{\alpha\beta} P_R \mu \right) \bar{N} \sigma_{\alpha\beta} N + h.c. \right] \quad .
 \end{aligned}$$

dipole

Effective Field Theory for $\mu \rightarrow e$ Conversion

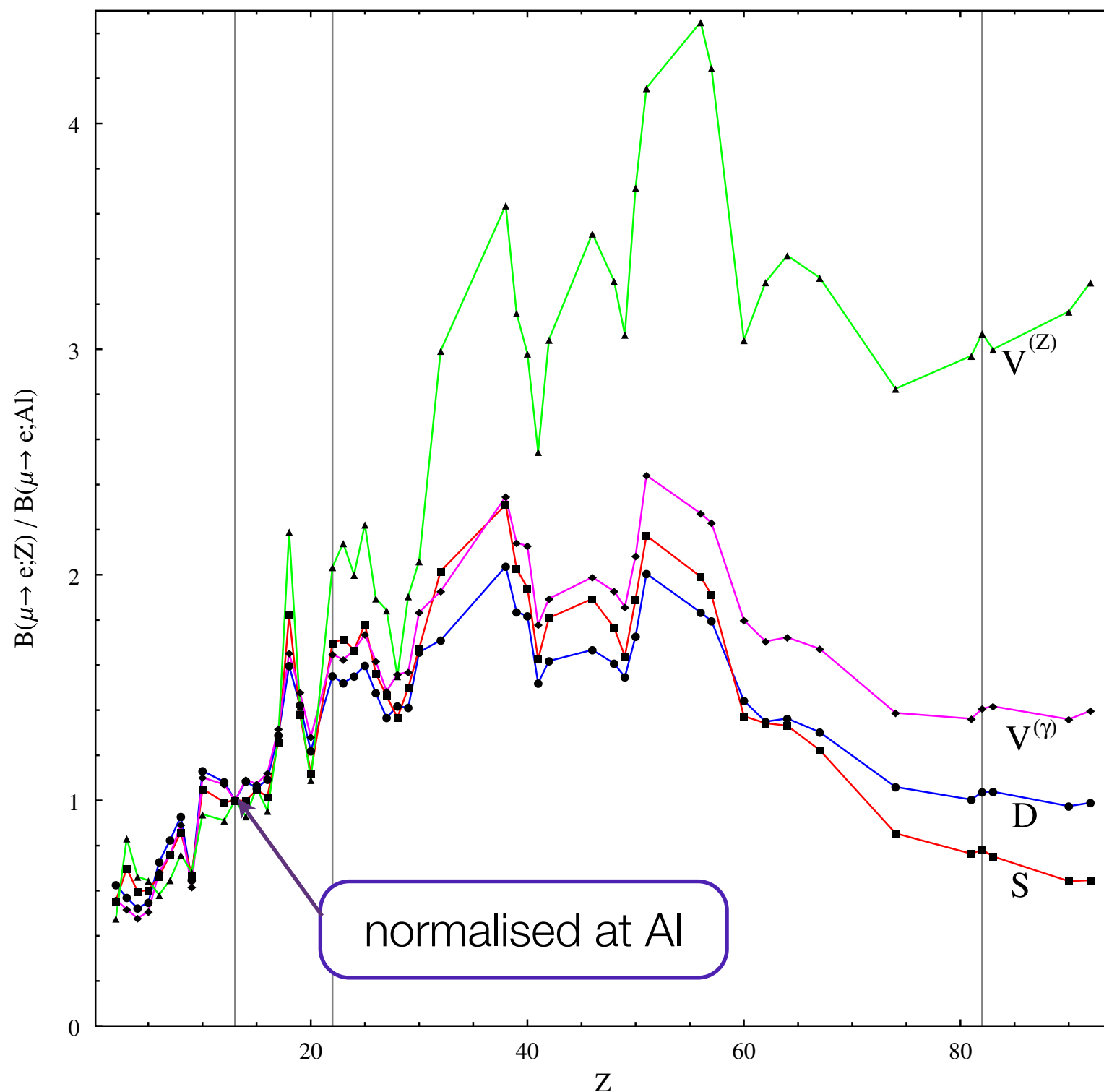


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 \end{aligned}$$

dipole

22 coeff. = 2 (dipole) + 2 (left/right) x 2 (proton/neutron) x 5 (interaction)

μ -e Conversion : Target dependence



V. Cirigliano, R. Kitano, Y. Okada,
and P. Tuzon, Phys. Rev. D80,
013002 (2009)

vector interaction
(with Z boson)

with Z
penguin

vector interaction
(with photon -
charge radius)

left-right
models

dipole interaction

SUSY-
GUT

scalar interaction

SUSY
seesaw

Spin Dependent μ -e conversion



Spin Dependent μ -e conversion



dipole
interaction

vector
interaction

scalar
interaction

Coherent
 μ -e Conversion
(spin independent)

Spin Dependent μ -e conversion



dipole
interaction

vector
interaction

scalar
interaction

Coherent
 μ -e Conversion
(spin independent)

axial vector
interaction

tensor
interaction

Incoherent
 μ -e Conversion
(spin dependent)

compare zero-spin and non-zero-spin nuclear targets

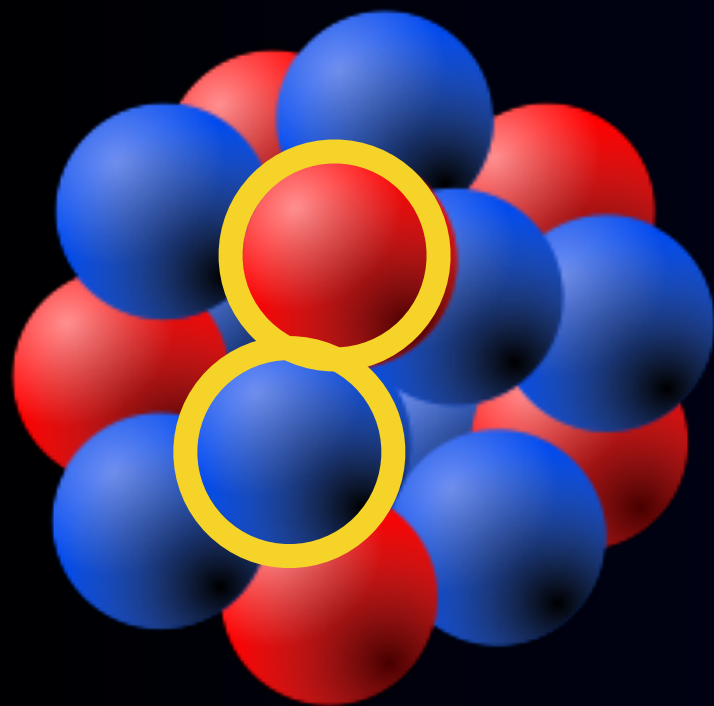
V. Cirigliano, S. Davidson, YK, Phys. Lett. B 771 (2017) 242

S. Davidson, YK, A. Saporta, Eur. Phys. J. C78 (2018) 109

Coherent and Incoherent : Pros and Cons

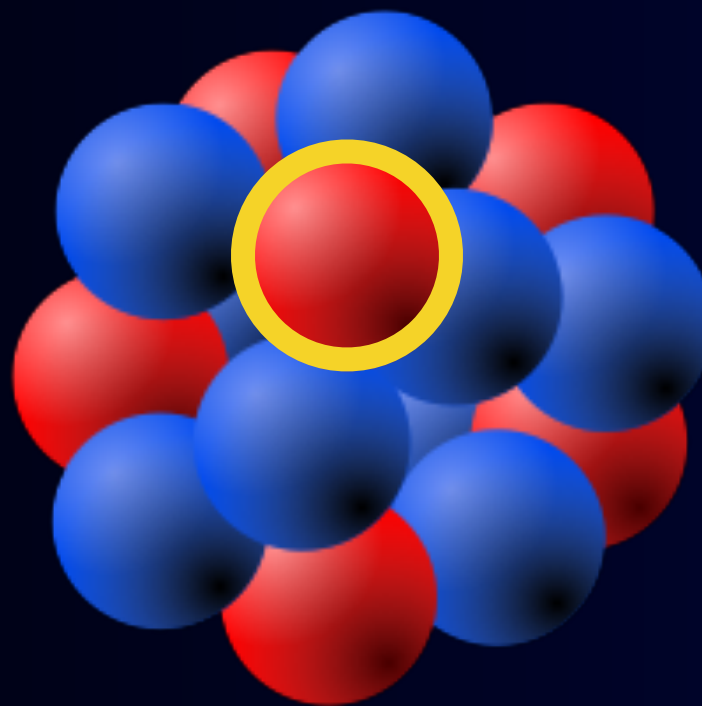


Coherent
μ-e Conversion
(spin independent)



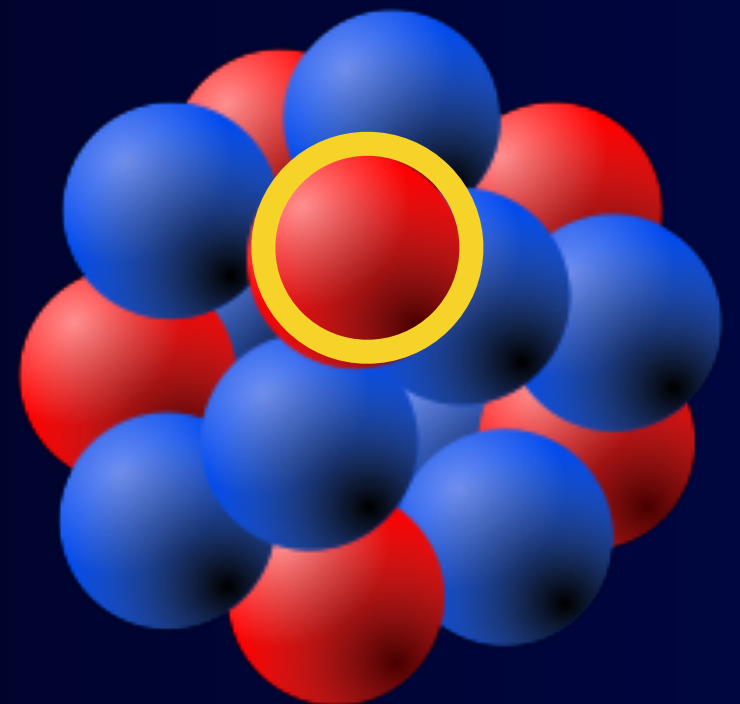
$$|\Sigma_i N_i|^2 \propto A^2$$

nuclear muon
capture



$$\Sigma_i |N_i|^2 \propto Z$$

Incoherent
μ-e Conversion
(spin dependent)



$$|N_i|^2 \propto 1$$

Other Muon LFV Processes

μ^- to e^+ conversion in muonic atom

$$\mu^- + N(A, Z) \rightarrow e^+ + N(A, Z - 2)$$

Lepton number violation (LNV) and
Lepton flavour violation (LFV)

Final can be the ground or excited states.

signal signature

$$E_{\mu e^+} = m_\mu - B_\mu - E_{rec} - (M(A, Z - 2) - M(A, Z))$$

backgrounds

- radiative muon nuclear capture (RMC)

$$\mu^- + N(A, Z) \rightarrow N(A, Z - 1) + \nu + \gamma$$

$$E_{RMC} = m_\mu - B_\mu - E_{rec} - (M(A, Z - 1) - M(A, Z))$$

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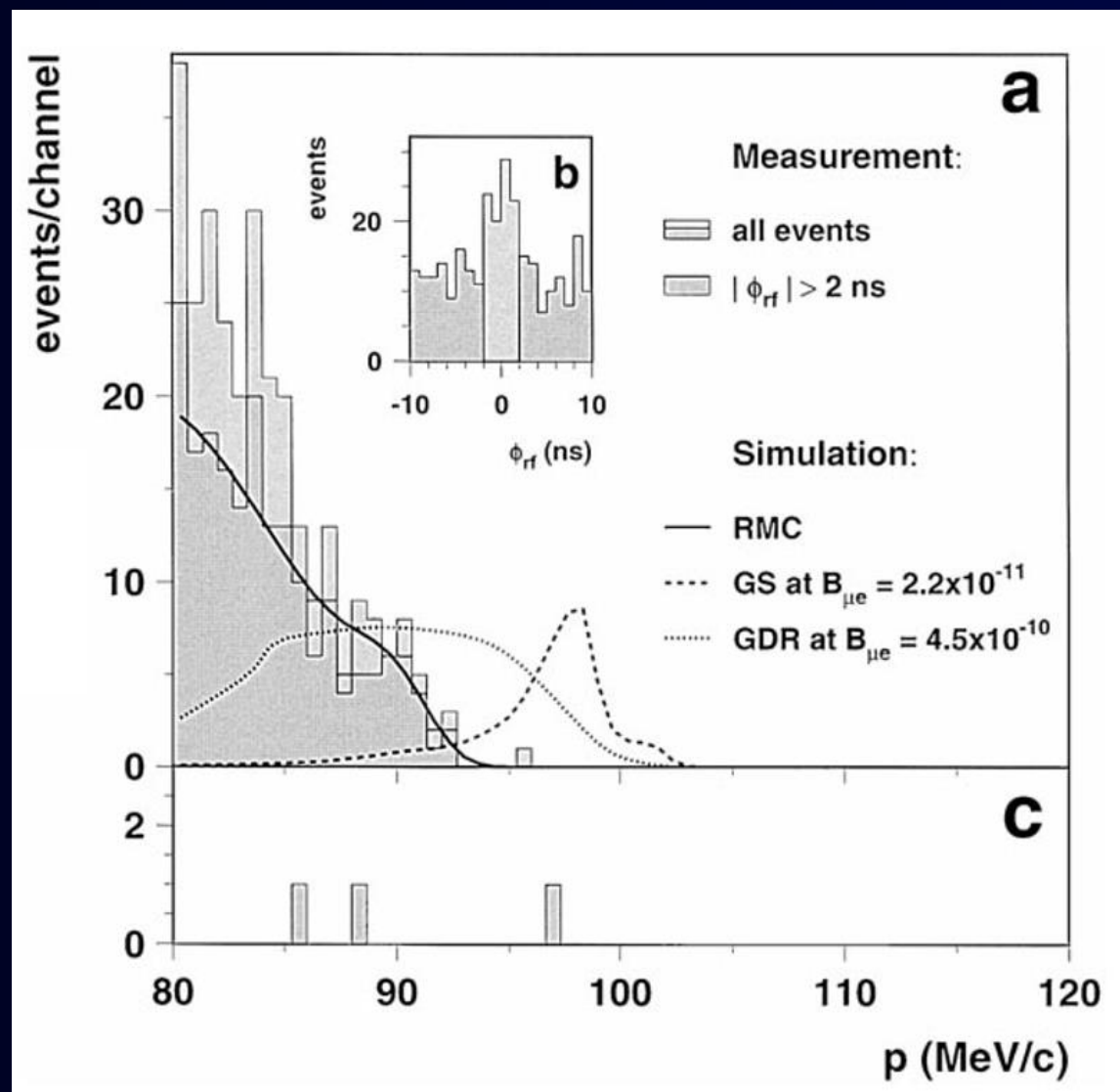
backgrounds

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$$\begin{array}{ll} \mu^- + \text{Ti} \rightarrow e^+ + \text{Ca}(\text{gs}) & 1.7 \times 10^{-12} \\ \mu^- + \text{Ti} \rightarrow e^+ + \text{Ca}(\text{ex}) & 3.6 \times 10^{-11} \end{array}$$



μ^- to e^+ conversion in the future



$$\mu^- + N(A, Z) \rightarrow e^+ + N(A, Z - 2)$$

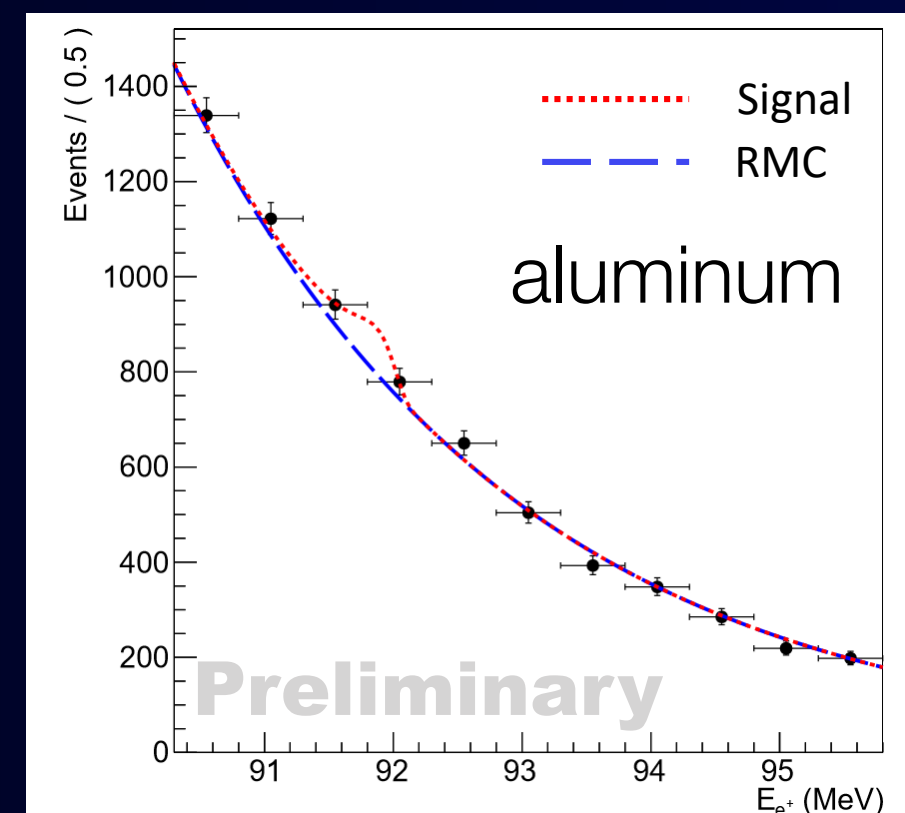
μ^- to e^+ conversion in the future

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Requirement on targets $E_{\mu e^+} > E_{RMC} \rightarrow M(A, Z - 1) < M(A, Z - 2)$

Atom	$E_{\mu^- e^+}$ (MeV)	$E_{\mu^- e^-}$ (MeV)	E_{RMC}^{end} (MeV)	N.A. (%)	f_{cap} (%)	τ_{μ^-} (ns)	A_T
^{27}Al	92.30	104.97	101.34	100	61.0	864	0.191
^{32}S	101.80	104.76	102.03	95.0	75.0	555	0.142
^{40}Ca	103.55	104.39	102.06	96.9	85.1	333	0.078
^{48}Ti	98.89	104.18	99.17	73.7	85.3	329	0.076
^{50}Cr	104.06	103.92	101.86	4.4	89.4	234	0.038
^{54}Fe	103.30	103.65	101.93	5.9	90.9	206	0.027
^{58}Ni	104.25	103.36	101.95	68.1	93.1	152	0.009
^{64}Zn	103.10	103.04	101.43	48.3	93.0	159	0.011
^{70}Ge	100.67	102.70	100.02	20.8	92.7	167	0.013

simulation for 10^{18} muons
signal $\sim 1 \times 10^{-12}$



Aluminum (for COMET & Mu2e) is not good.

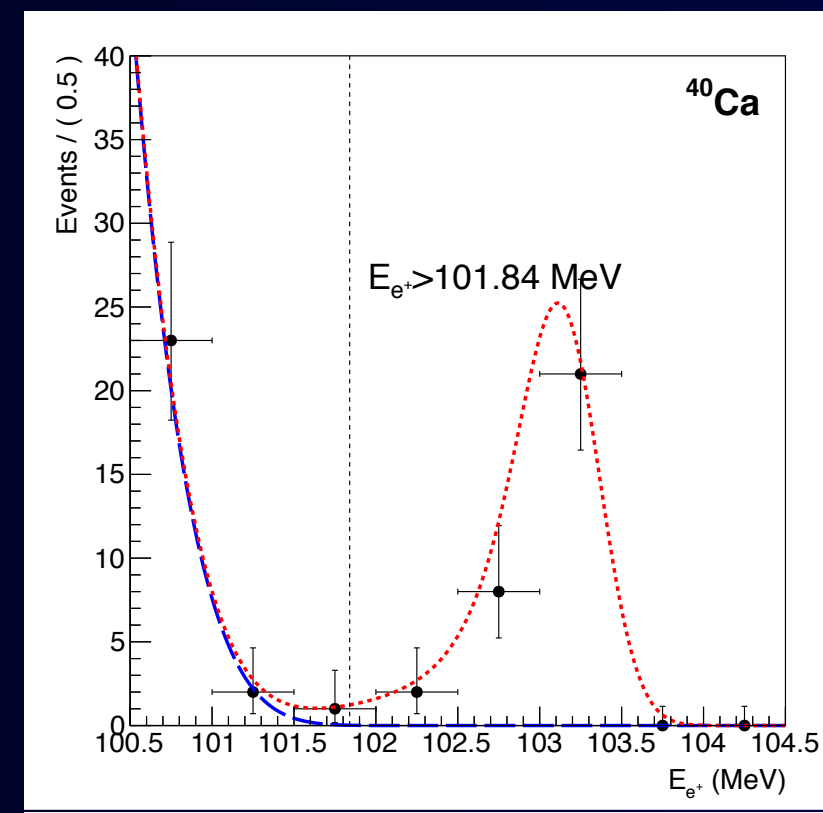
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signal $\sim 1 \times 10^{-14}$



Muonium to Antimuonium Conversion

$\text{Mu } (\mu^+e^-) \rightarrow \text{Mu } (\mu^-e^+)$



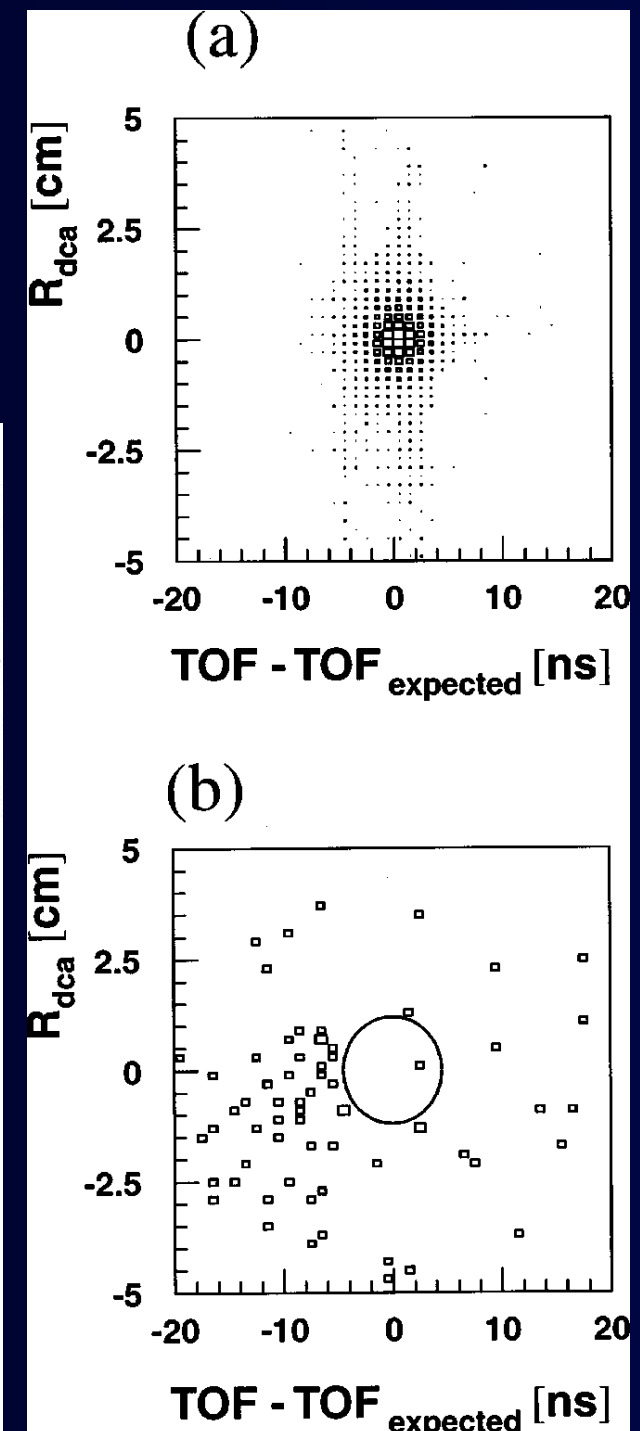
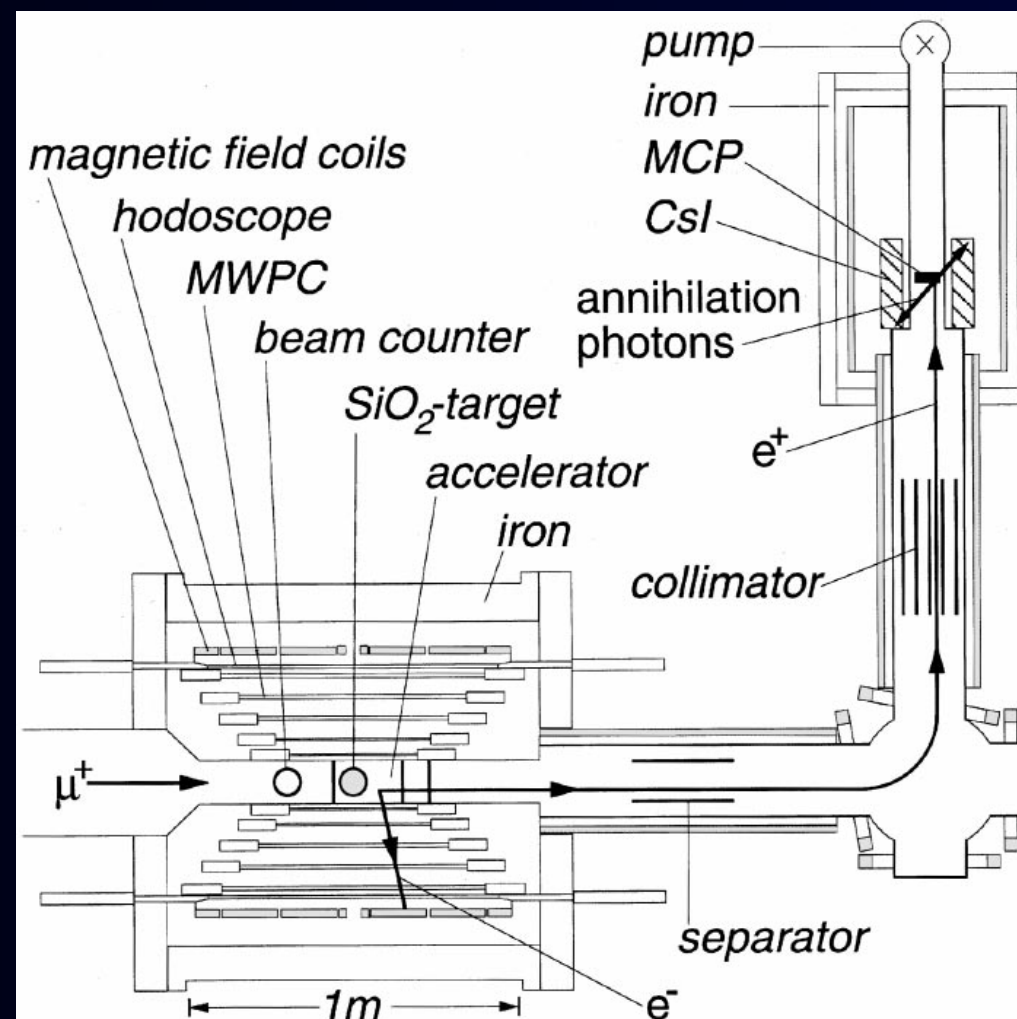
$$\mu^+ + e^- \rightarrow \mu^- + e^+$$

muonium to anti-muonium conversion

$$|\Delta L_{\mu/e}| = 2$$

- doubly-charged Higgs model etc.
- muonium production in vacuum
- previous experiment at PSI (1999)

$$G_{\text{Mu}\overline{\text{Mu}}} < 3 \times 10^{-3} G_F$$



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$$\mu^+ + e^- \rightarrow \mu^- + e^+$$

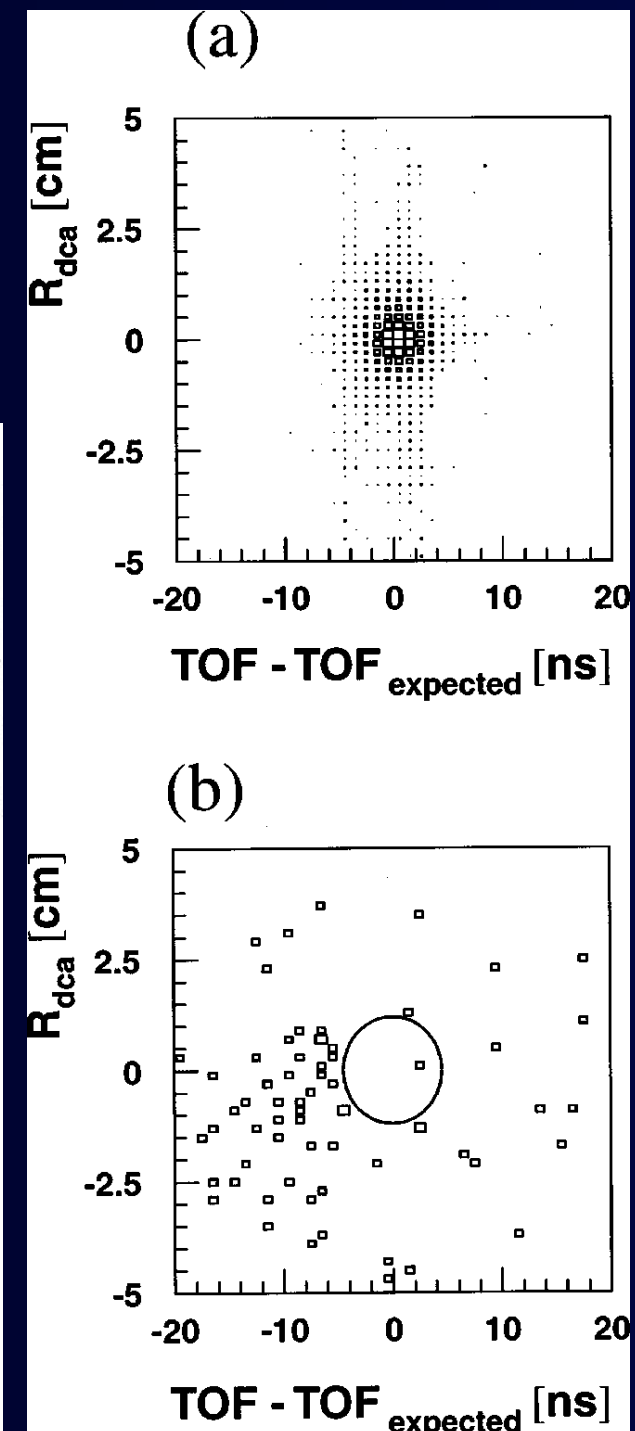
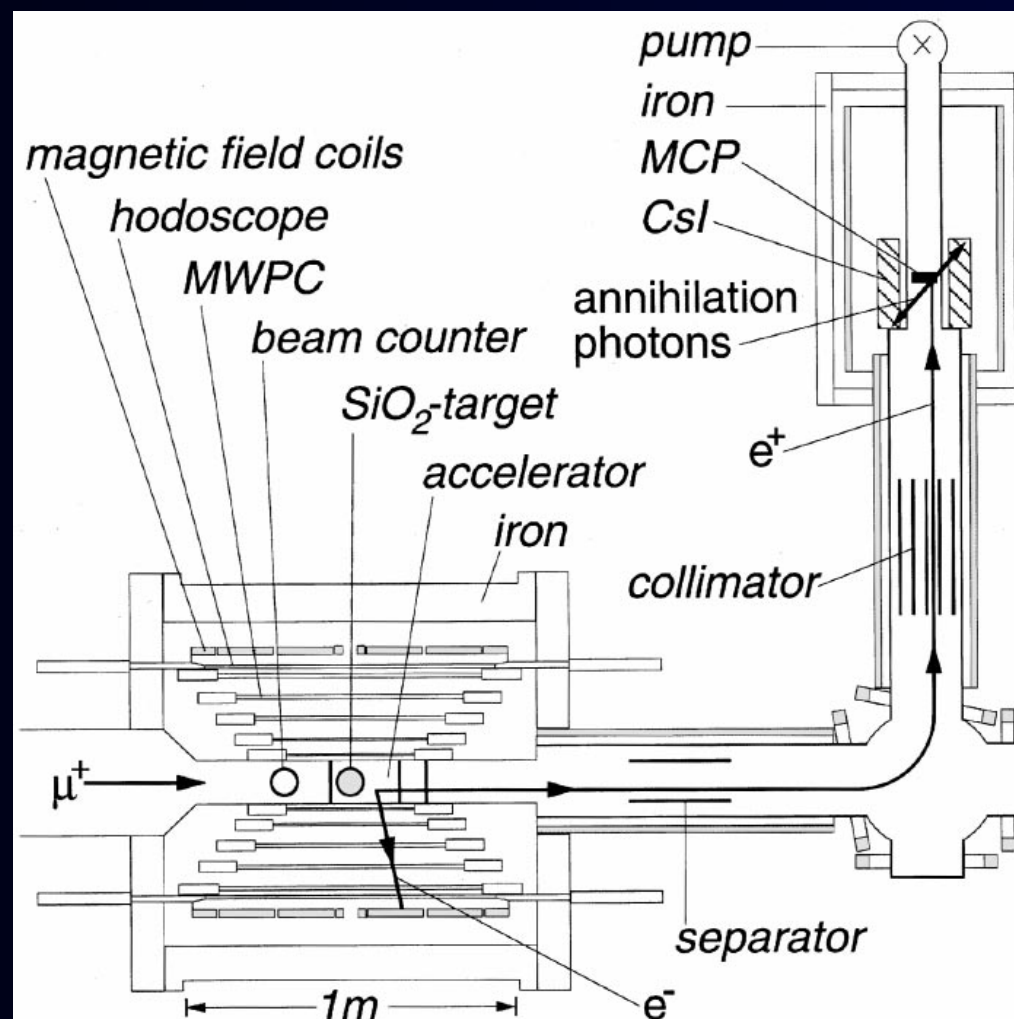
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new attempt at J-PARC?

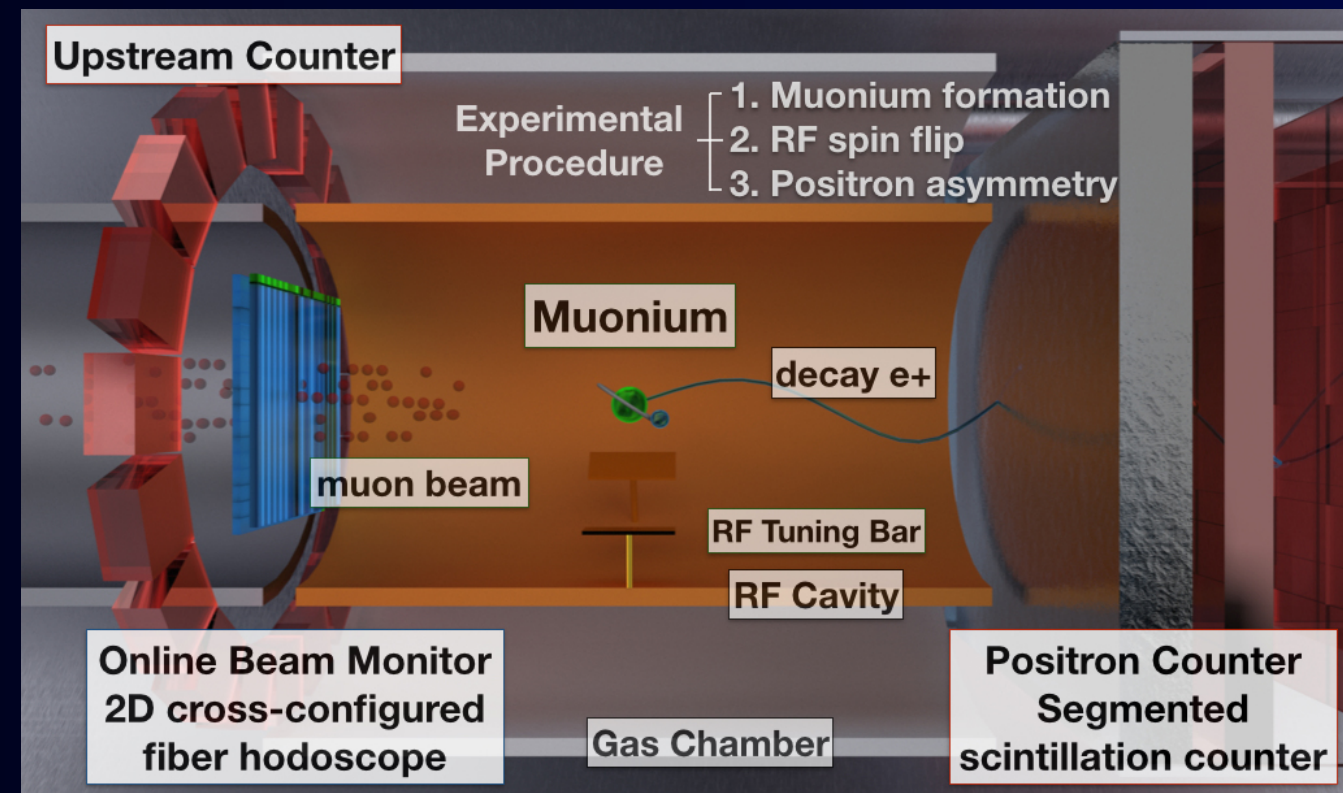


Muonium :



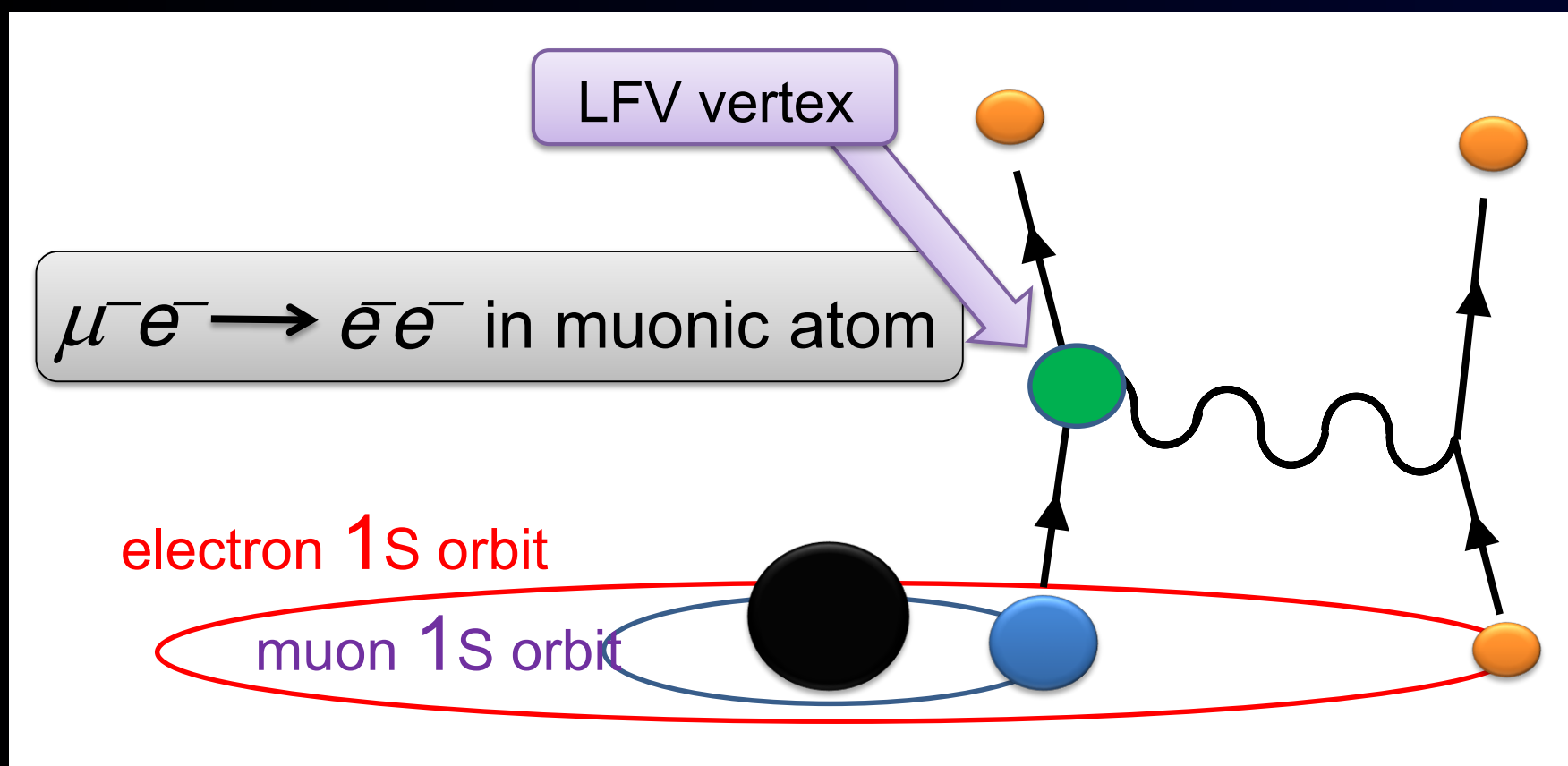
- similar to $\mu \rightarrow eee$
 - may be useful to distinguish different couplings
 - 2 body final state
- disadvantage
 - poor wave function overlap between μ and e
 - Coulomb bound state
- no experiments so far
- muonium production in vacuum at Muse @ J-PARC
 - 10^{15} for 2×10^7 sec
 - $< 10^{-13}$ might be possible in the near future.

Museum detector
@J-PARC



$\mu^- + e^- \rightarrow e^- + e^-$ in a muonic atom

$$\mu^- + e^- \rightarrow e^- + e^-$$



$\mu^- e^- \rightarrow e^- e^-$ has the overlap of μ^- and e^- which is proportional to Z^3 . (almost compatible to $\mu^+ \rightarrow e^+ e^+ e^-$)

Experimentally a pair of e^- and e^- in the final state is measured.

Z dependence discriminate dipole and contact contributions.

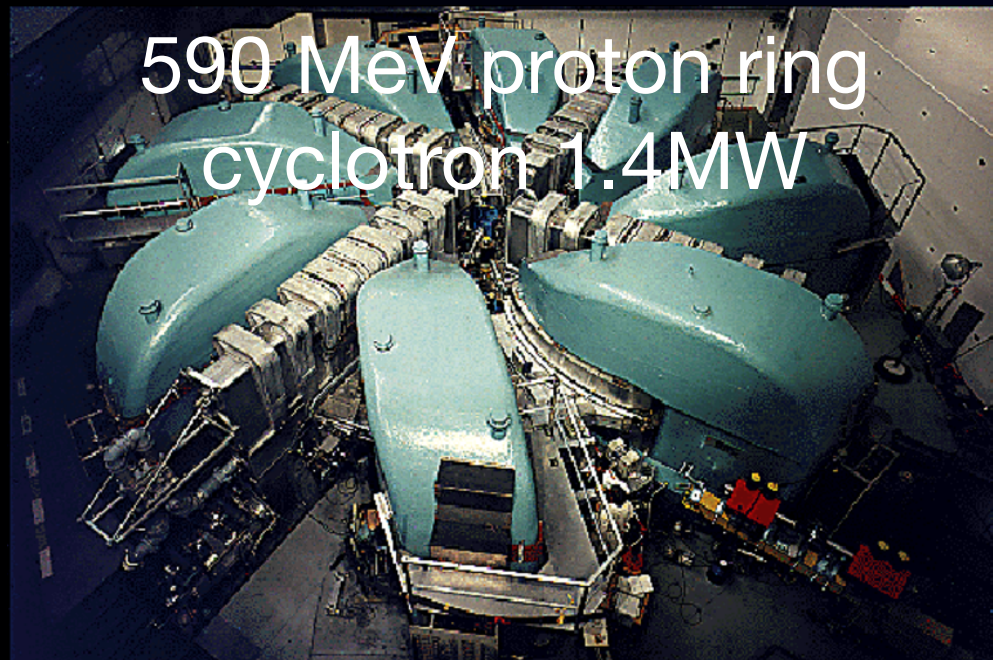
M. Koike, YK, J. Sato and M. Yamanaka, Phys. Rev. Lett. 105 (2010)

Y. Uesaka, YK, J. Sato, T. Sato and M. Yamanaka, Phys. Rev. D93 (2016) 076006

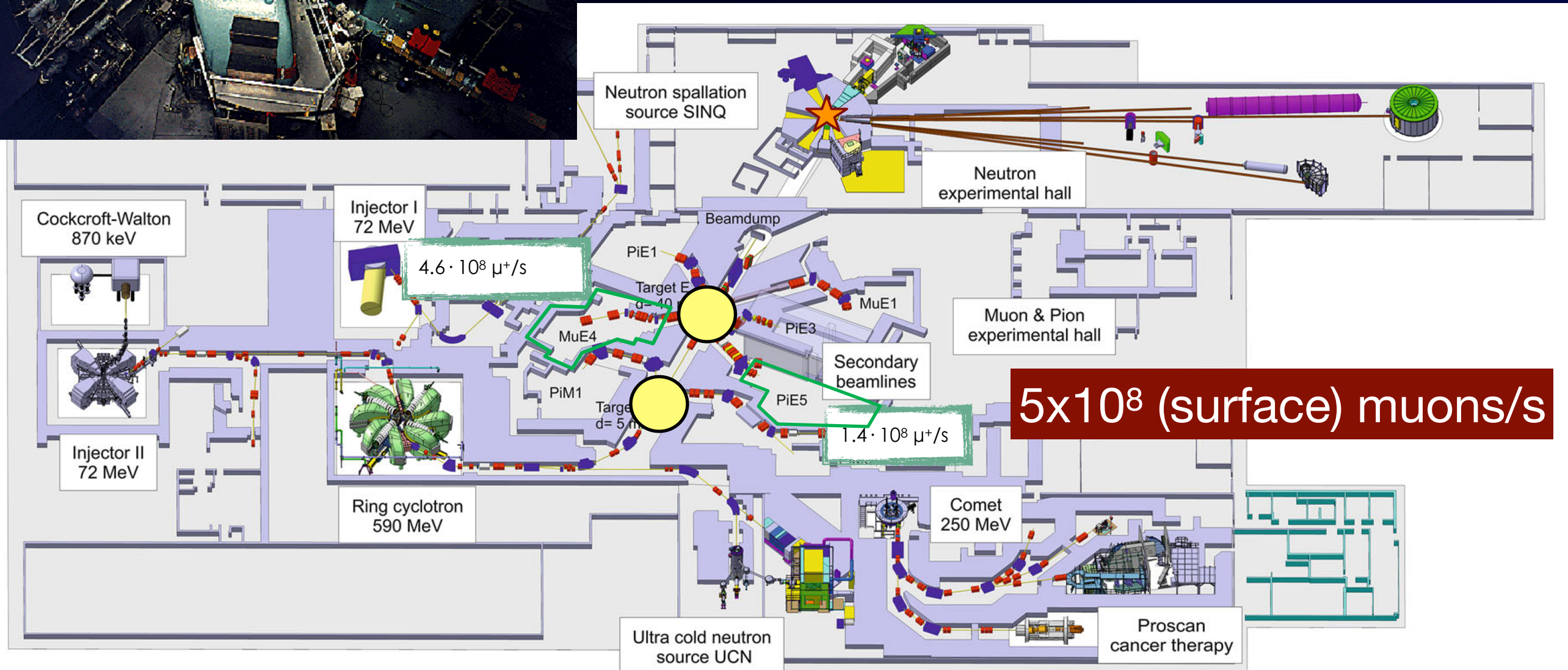
Y. Uesaka, YK, J. Sato, T. Sato and M. Yamanaka, Phys. Rev. D97 (2018) 015017

Muon Facilities for CLFV

Paul Sherrer Institute (PSI)



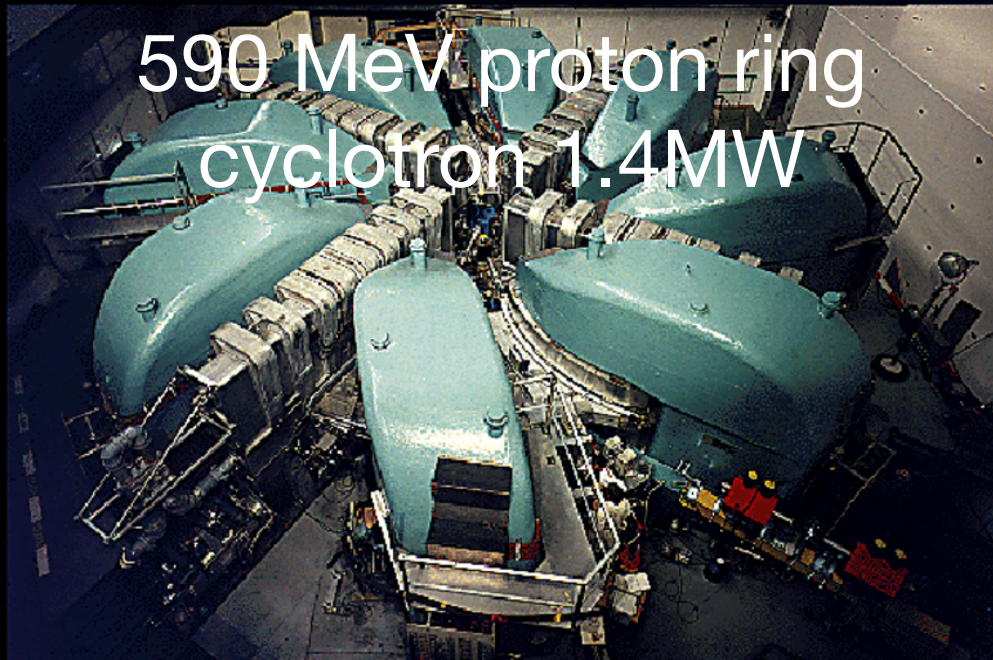
590 MeV proton ring
cyclotron 1.4MW



Paul Sherrer Institute (PSI)



590 MeV proton ring
cyclotron 1.4 MW



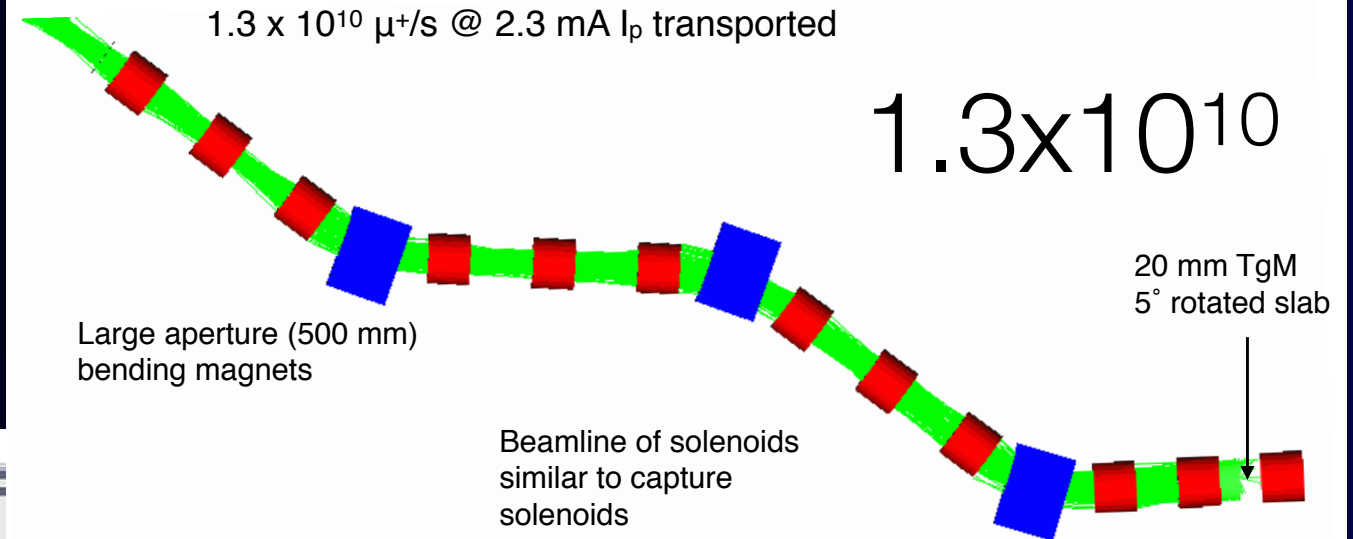
$1.3 \times 10^{10} \mu^+/s$ @ 2.3 mA I_p transported

1.3×10^{10}

Large aperture (500 mm)
bending magnets

Beamline of solenoids
similar to capture
solenoids

20 mm TgM
5° rotated slab



Neutron spallation
source SINQ

Neutron
experimental hall

Cockcroft-Walton
870 keV

Injector I
72 MeV

$4.6 \cdot 10^8 \mu^+/s$

Beamdump

PIE1

Target E

$\theta = 40^\circ$

MuE4

MuE1

PIE3

Muon & Pion
experimental hall

Secondary
beamlines

Target
 $d = 5$ m

PIE5

$1.4 \cdot 10^8 \mu^+/s$

5×10^8 (surface) muons/s

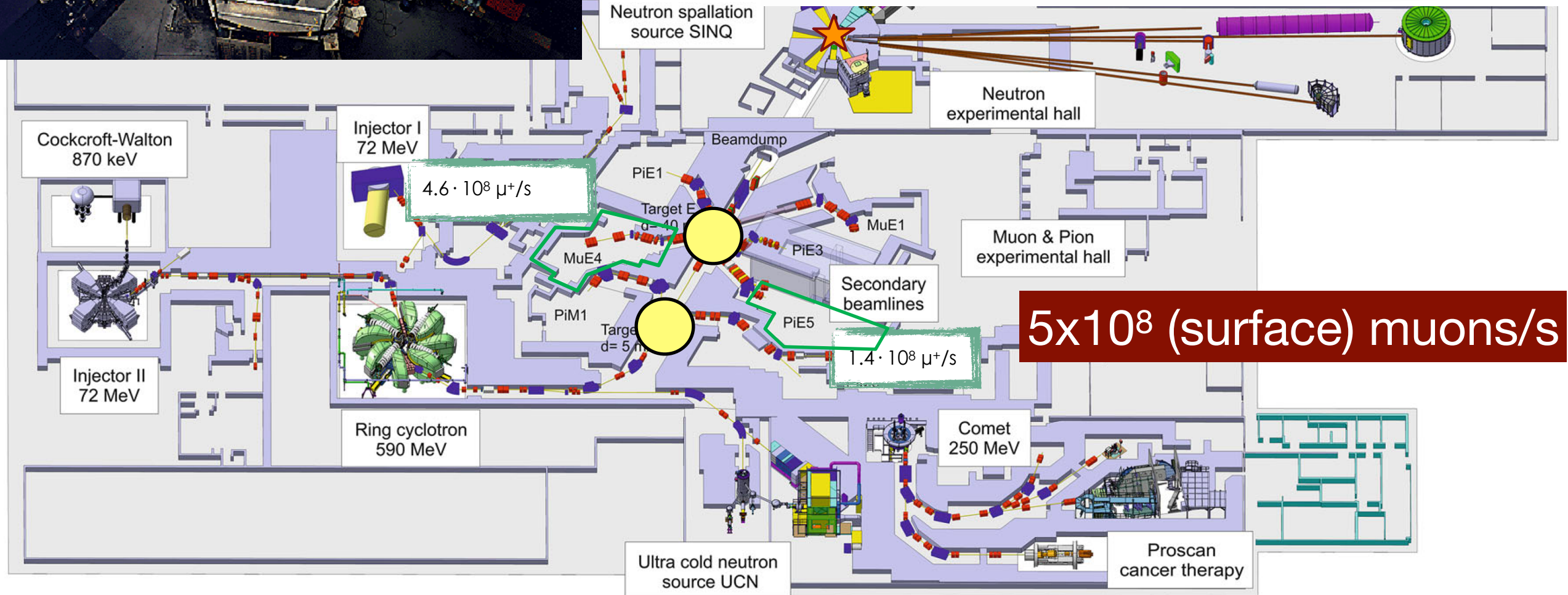
Injector II
72 MeV

Ring cyclotron
590 MeV

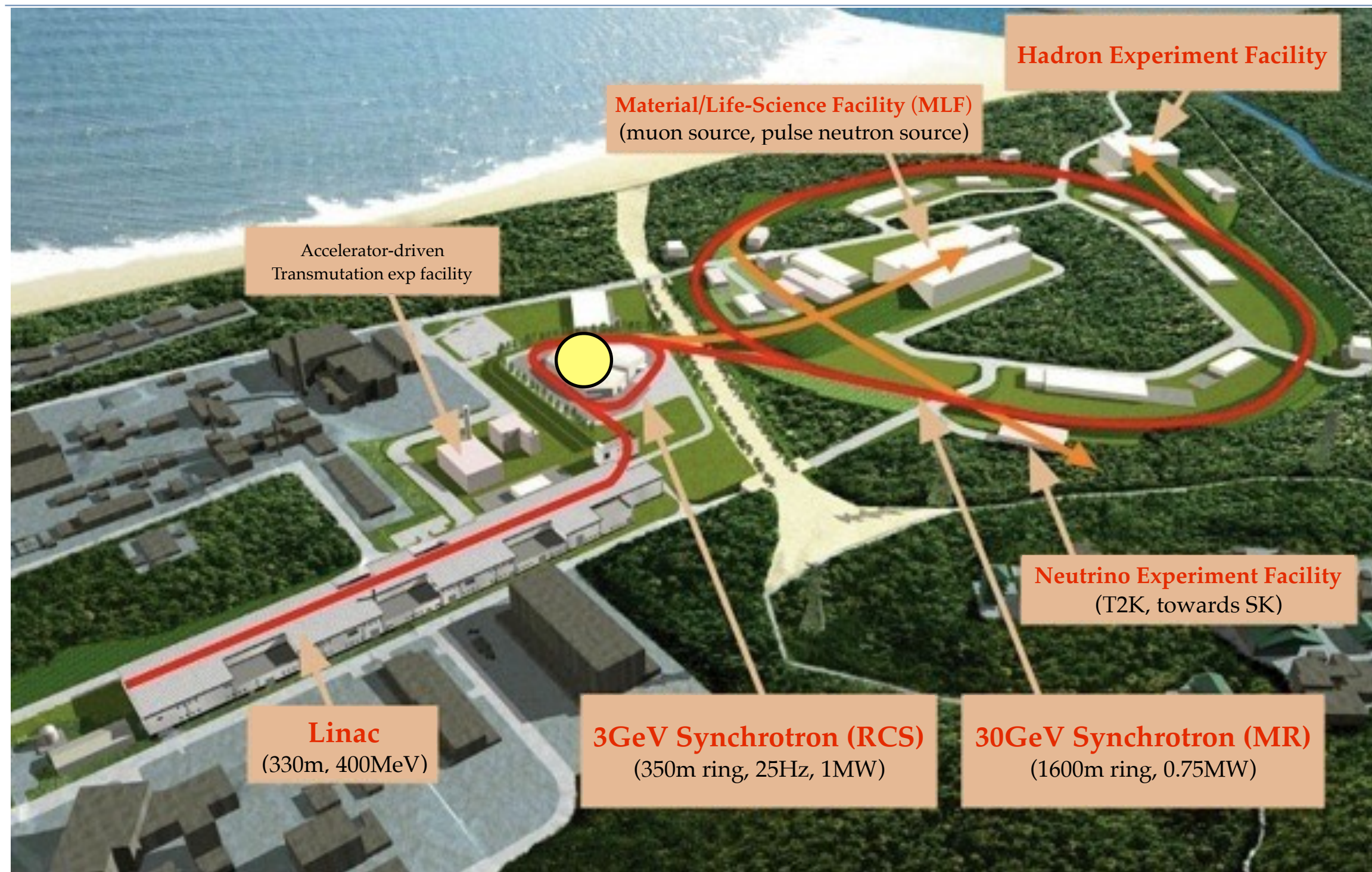
Comet
250 MeV

Ultra cold neutron
source UCN

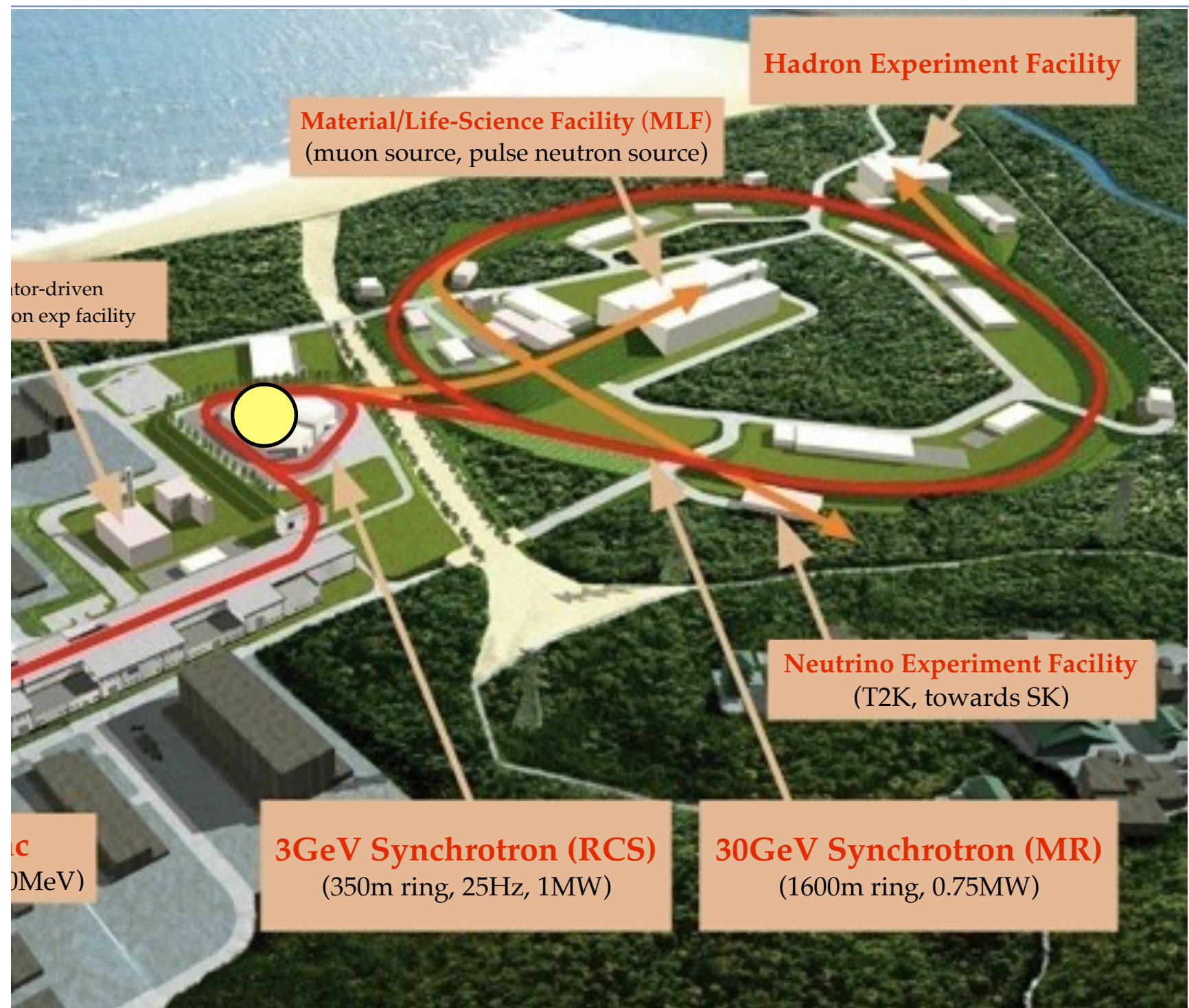
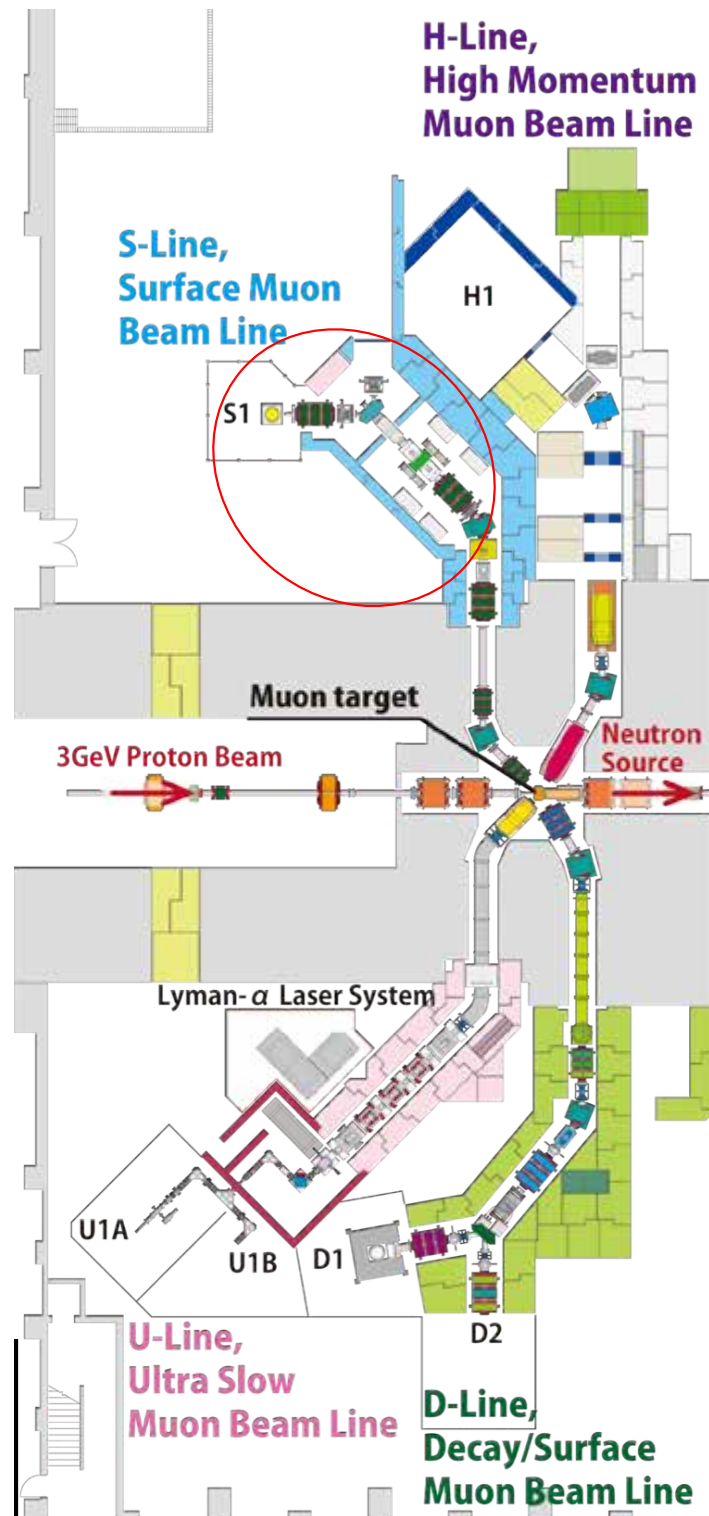
Proscan
cancer therapy



J-PARC (Muse@MLF)



J-PARC (Muse@MLF)



MuSIC at RCNP, Osaka University

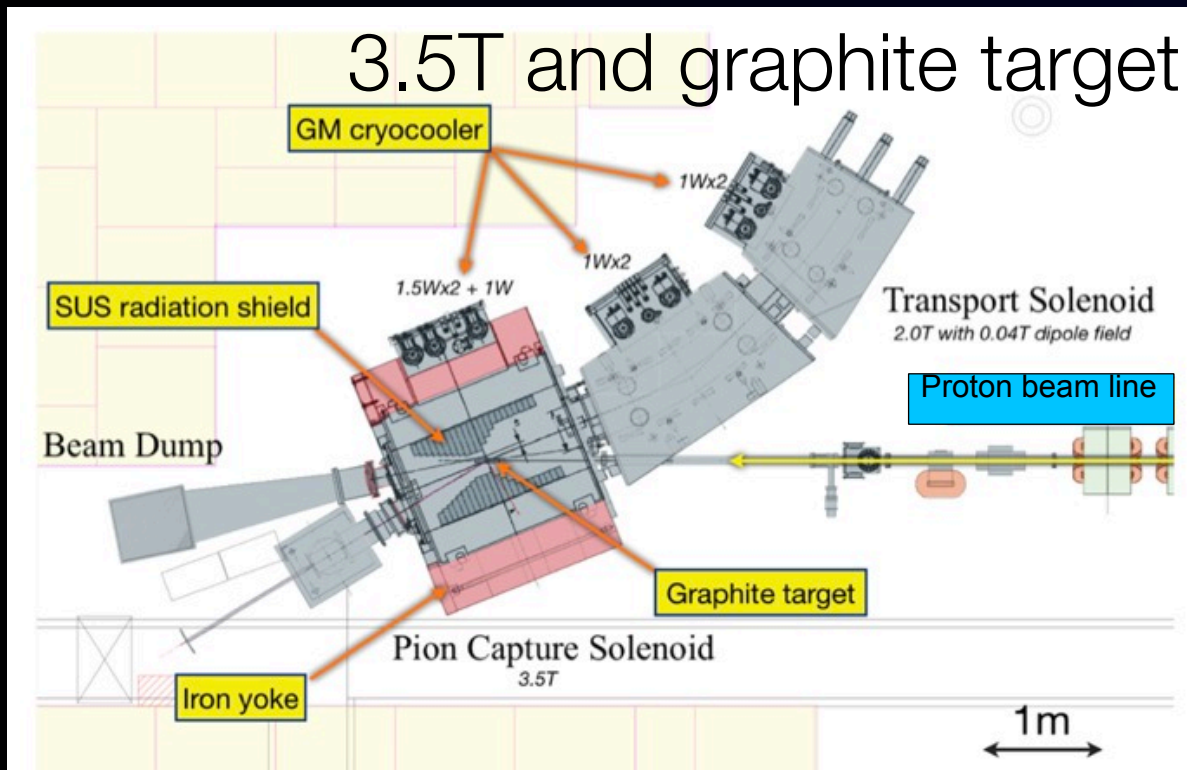


3.5T and graphite target

MuSIC at RCNP, Osaka University



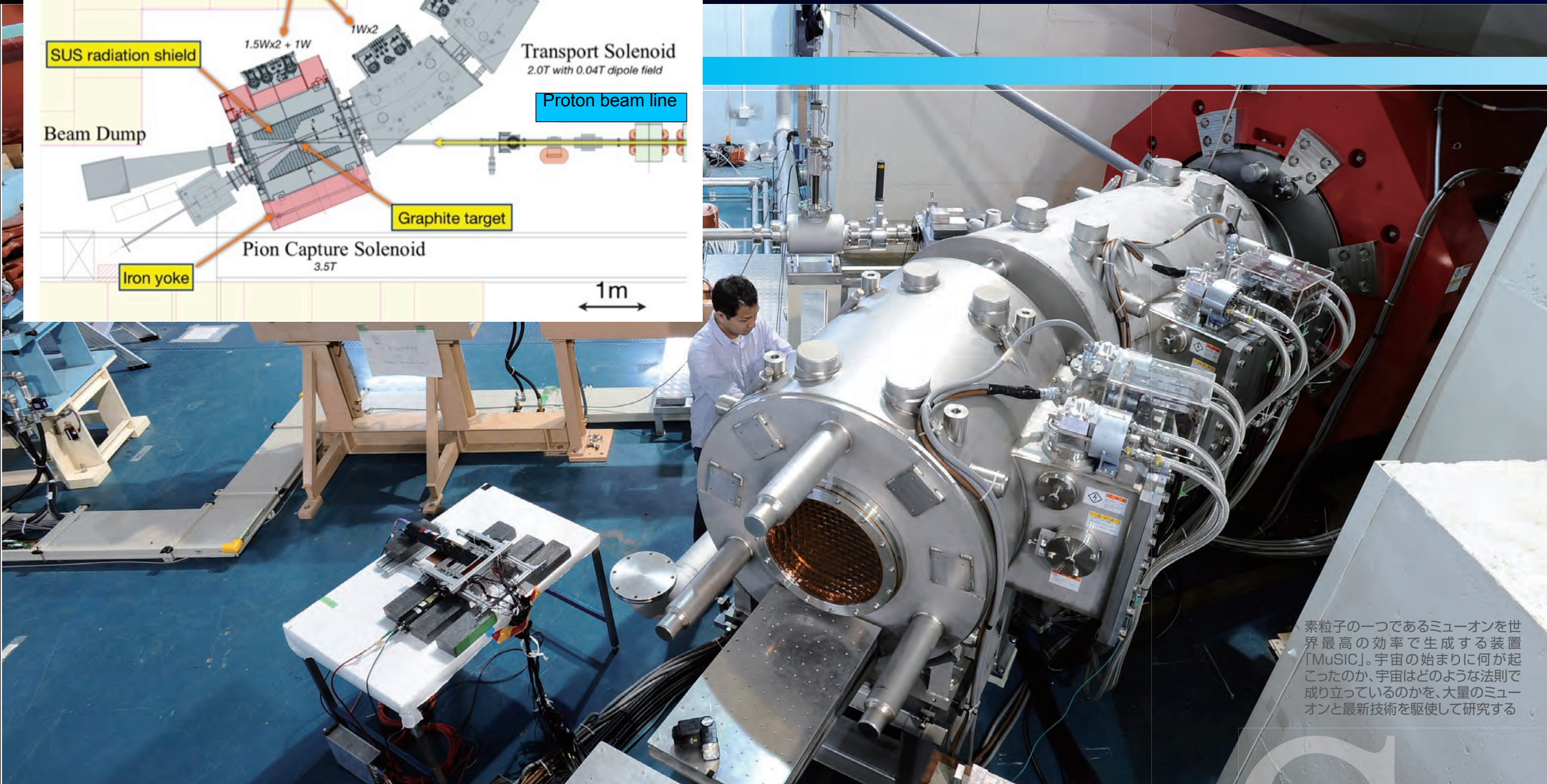
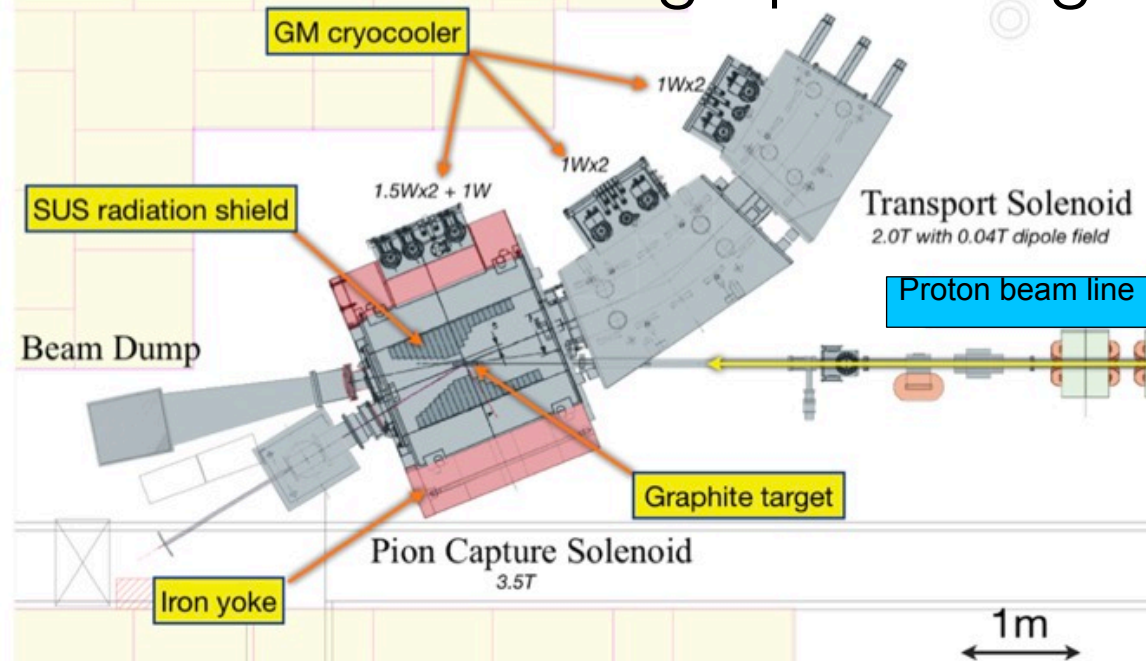
3.5T and graphite target



MuSIC at RCNP, Osaka University



3.5T and graphite target

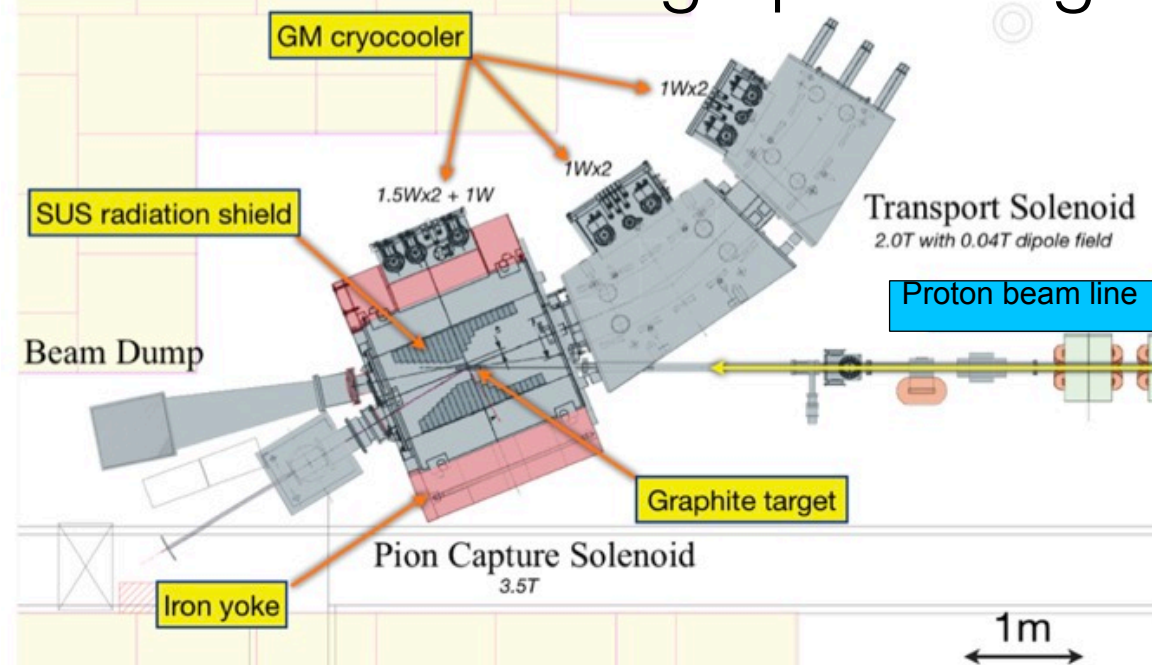


素粒子の一つであるミューオンを世界最高の効率で生成する装置「MuSIC」。宇宙の始まりに何が起こったのか、宇宙はどのような法則で成り立っているのかを、大量のミューオンと最新技術を駆使して研究する

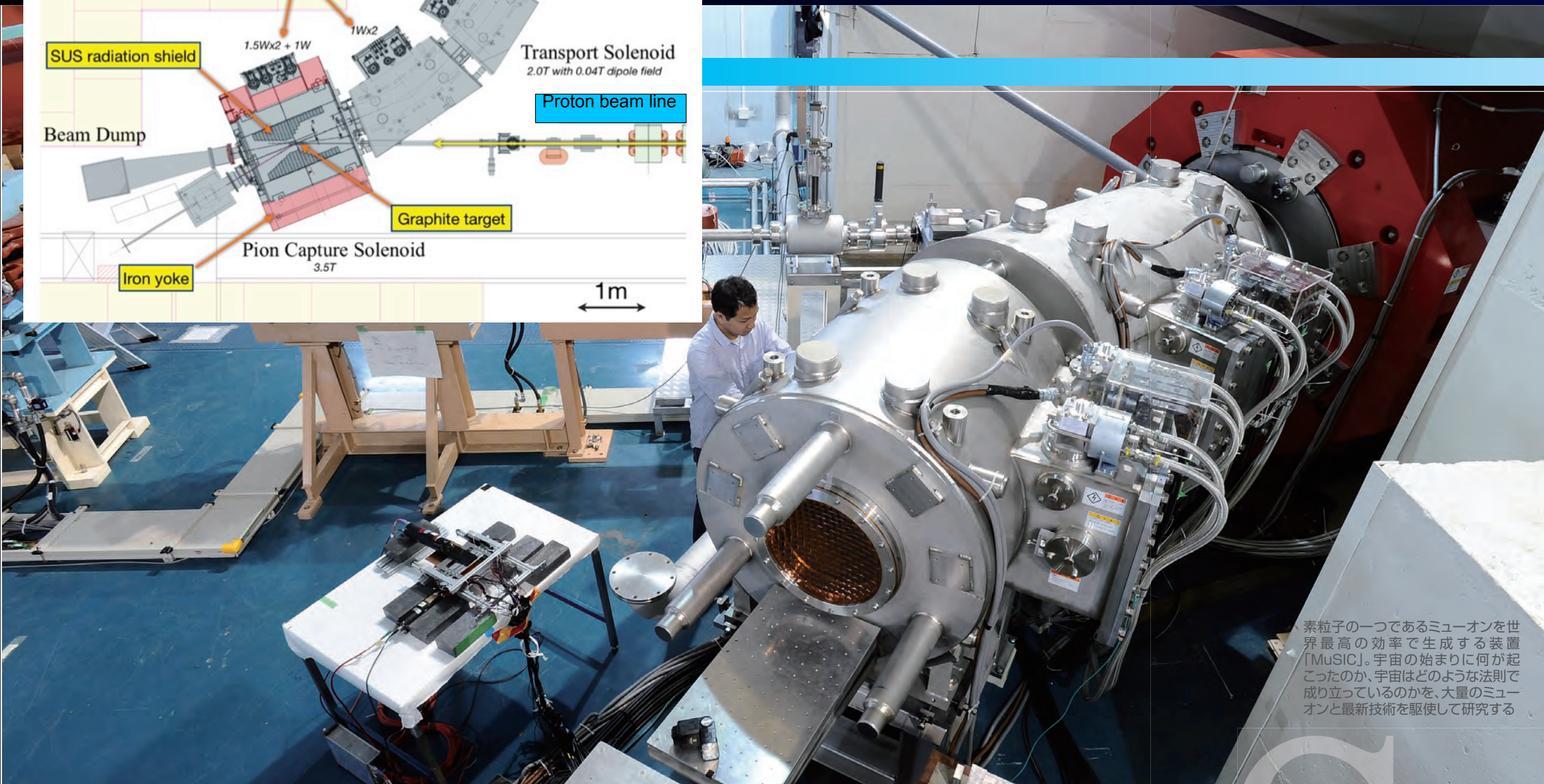
MuSIC at RCNP, Osaka University



3.5T and graphite target



Muon Science Intense Channel (>2011)

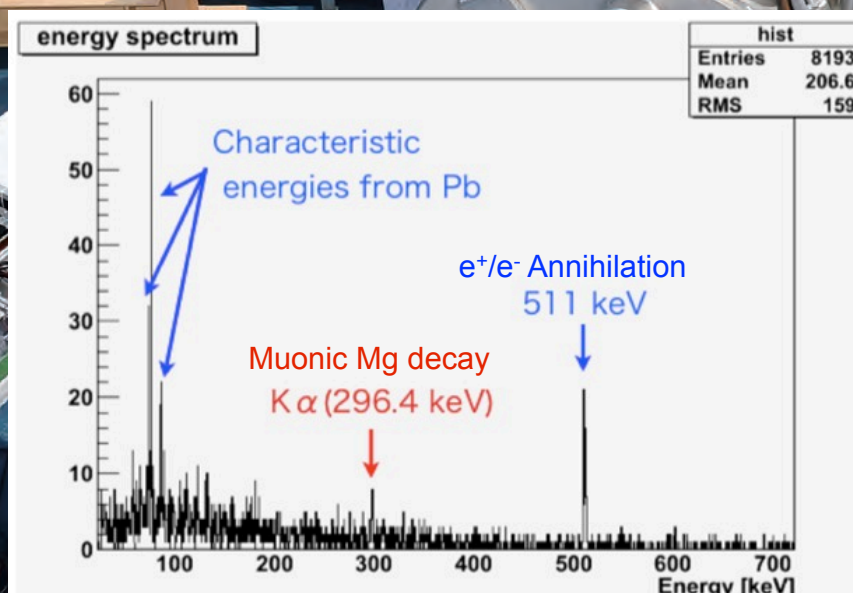
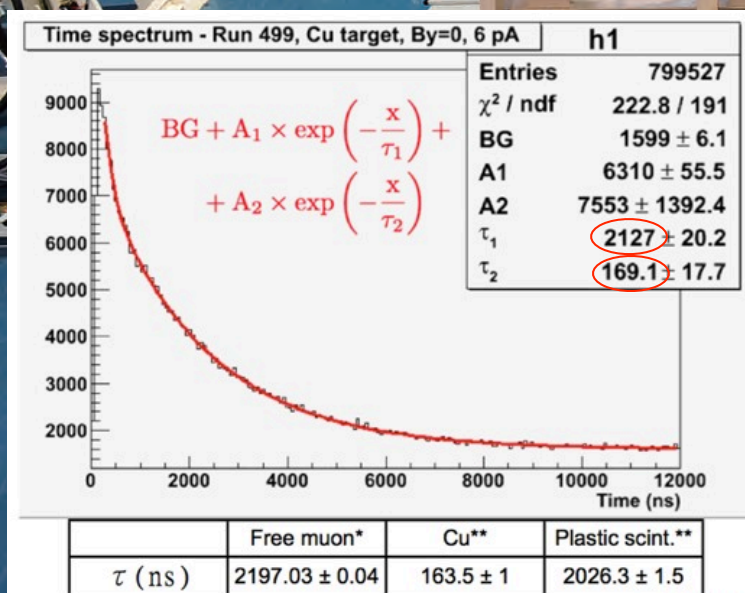
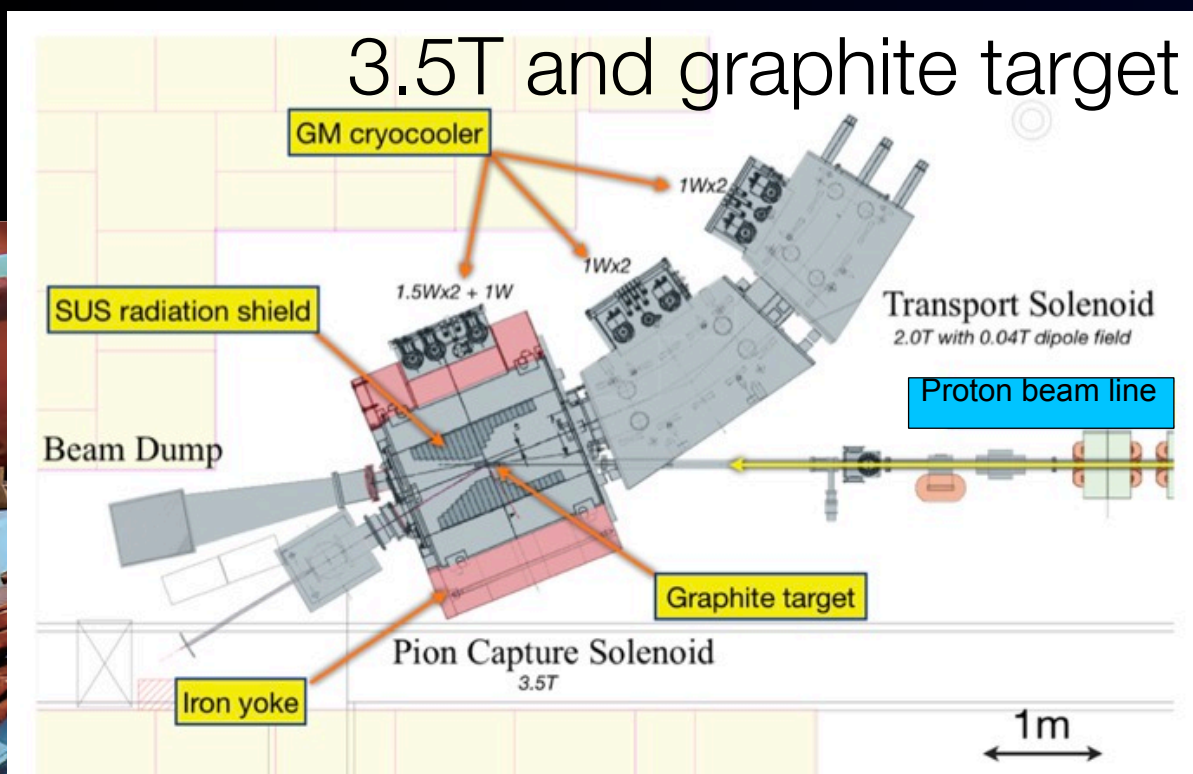


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MuSIC at RCNP, Osaka University



Muon Science Intense Channel (>2011)



MuSIC muon yields

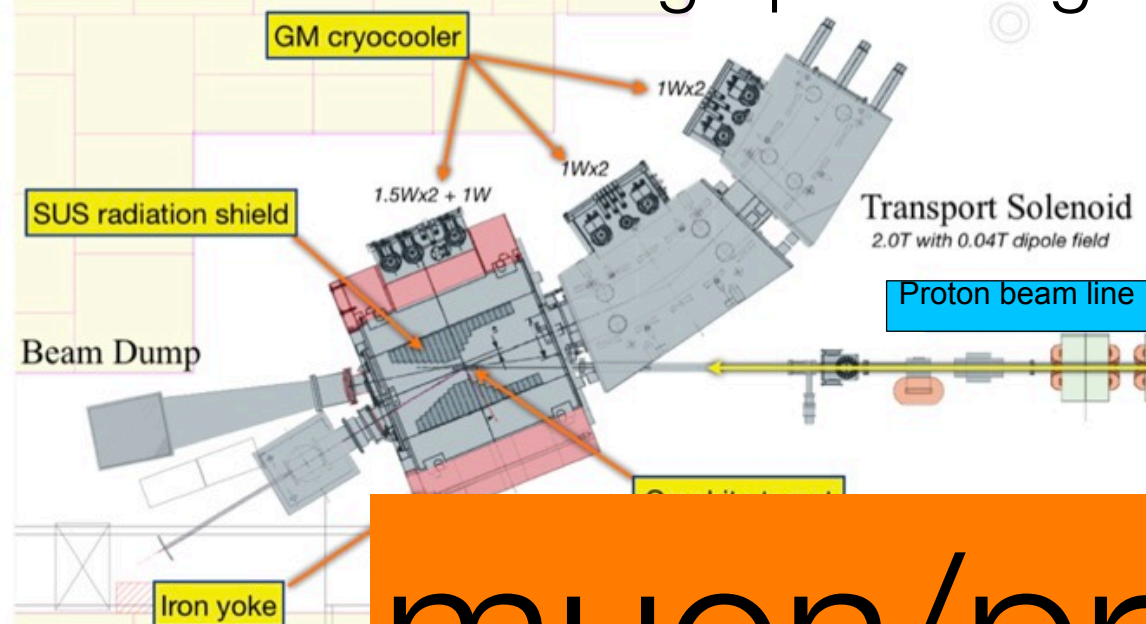
μ^+ : $3 \times 10^8/\text{s}$ for 400W

μ^- : $1 \times 10^8/\text{s}$ for 400W

MuSIC at RCNP, Osaka University

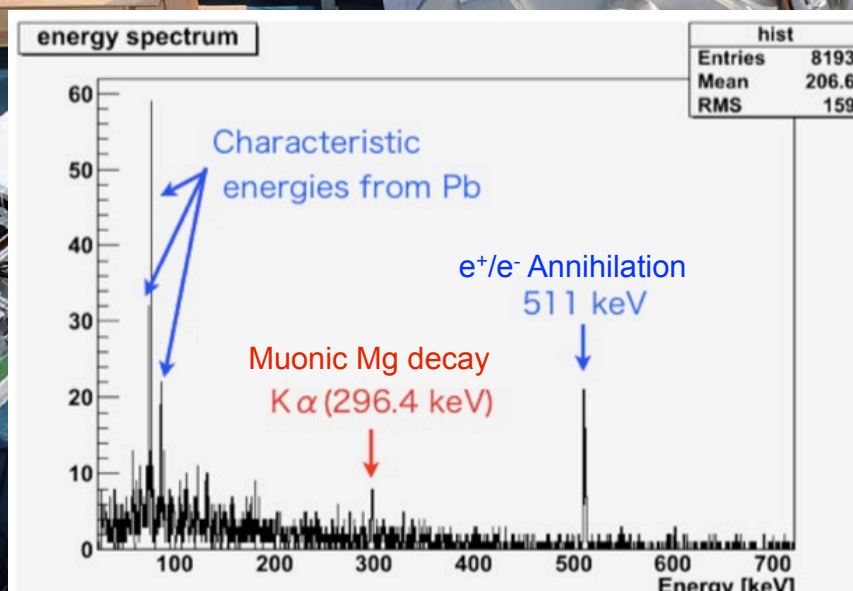
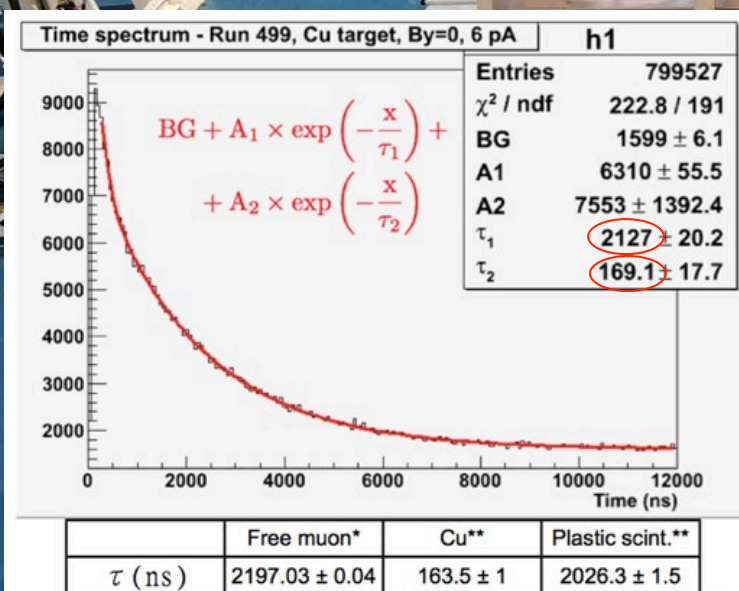


3.5T and graphite target



Muon Science Intense Channel (>2011)

muon/proton ~ x1000



MuSIC muon yields

μ^+ : $3 \times 10^8/\text{s}$ for 400W

μ^- : $1 \times 10^8/\text{s}$ for 400W



Delivering the world's most intense muon beam

S. Cook,¹ R. D'Arcy,¹ A. Edmonds,¹ M. Fukuda,² K. Hatanaka,² Y. Hino,³ Y. Kuno,³
M. Lancaster,¹ Y. Mori,⁴ T. Ogitsu,⁵ H. Sakamoto,³ A. Sato,³ N. H. Tran,³ N. M. Truong,³
M. Wing,^{1,*} A. Yamamoto,⁵ and M. Yoshida⁵

¹*Department of Physics and Astronomy, UCL, Gower Street, London WC1E 6BT, United Kingdom*

²*Research Center for Nuclear Physics (RCNP), Osaka University, Osaka 567-0047, Japan*

³*Department of Physics, Graduate School of Science, Osaka University, Osaka 569-0043, Japan*

⁴*Kyoto University Reactor Research Institute (KURRI), Kyoto 590-0494, Japan*

⁵*High Energy Accelerator Research Organization (KEK), Tsukuba 305-0801, Japan*

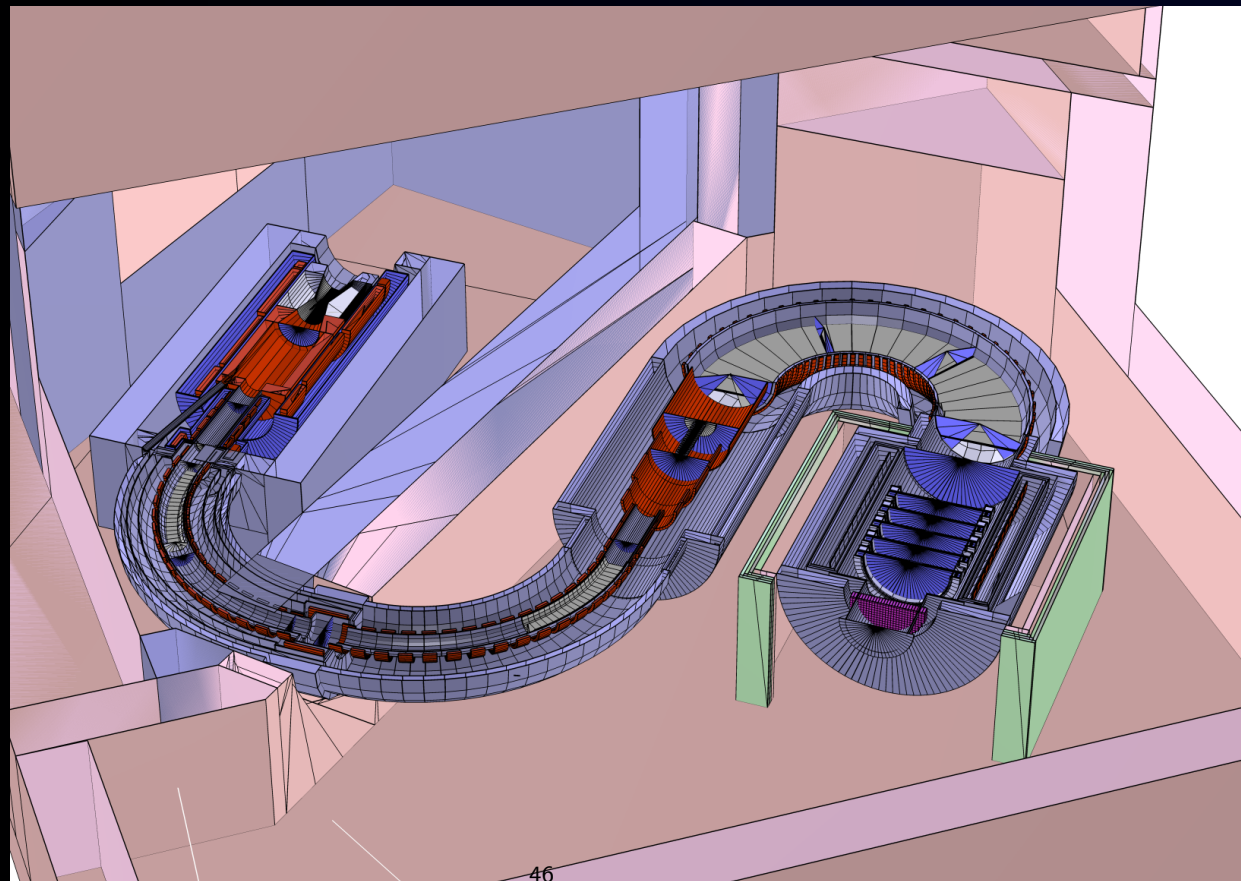
(Received 25 October 2016; published 15 March 2017)

A new muon beam line, the muon science innovative channel, was set up at the Research Center for Nuclear Physics, Osaka University, in Osaka, Japan, using the 392 MeV proton beam impinging on a target. The production of an intense muon beam relies on the efficient capture of pions, which subsequently decay to muons, using a novel superconducting solenoid magnet system. After the pion-capture solenoid, the first 36° of the curved muon transport line was commissioned and the muon flux was measured. In order to detect muons, a target of either copper or magnesium was placed to stop muons at the end of the muon

COMET Phase-I



COMET Phase-I

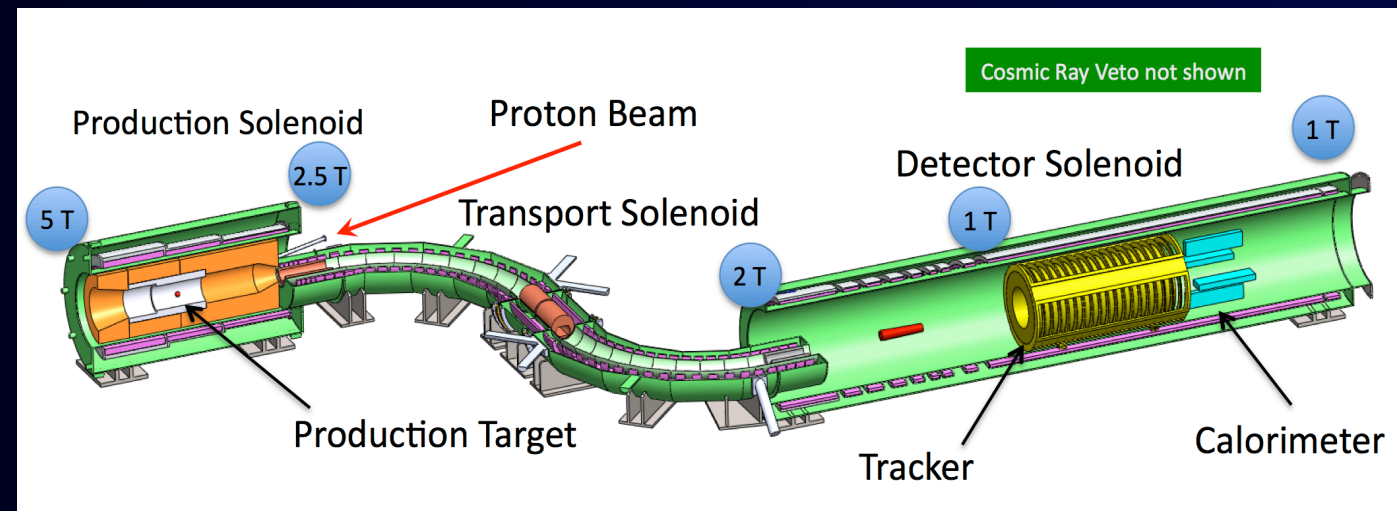


proton beam power=56 kW

$\sim 2.0 \times 10^{11}$ stopped muon/s

(0.004 stopped muon/proton)

Mu2e



proton beam power=8 kW

$\sim 1.2 \times 10^{10}$ stopped muon/s

stopped muons
 $\sim O(4 \times 10^{18})/\text{year}$

$\mu - e$ conversion

COMET@J-PARC

COMET Phase-I



proton beam power=3.2 kW

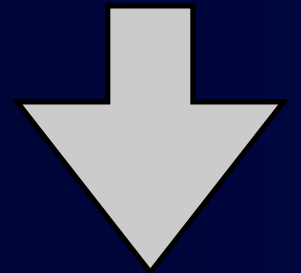
proton target

muon target

muon beamline

detector

current
limit:
 $< 6 \times 10^{-13}$



aimed
sensitivity:
 $< 6 \times 10^{-15}$

$\mathcal{O}(10^{-15})$

COMET Phase-I



proton beam power=3.2 kW

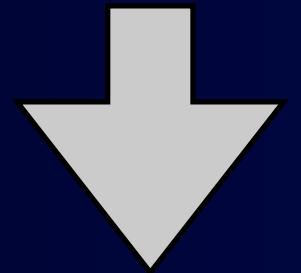
proton target

muon target

muon beamline

detector

current
limit:
 $< 6 \times 10^{-13}$



aimed
sensitivity:
 $< 6 \times 10^{-15}$

$\mathcal{O}(10^{-15})$

x100 improvement

COMET Phase-II



proton beam power=56 kW

proton target

electron spectrometer

muon target

muon beamline

detector

original
sensitivity:
 2.6×10^{-17}

$\mathcal{O}(10^{-17})$

COMET Phase-II



proton beam power=56 kW

proton target

electron spectrometer

muon target

muon beamline

detector

original
sensitivity:
 2.6×10^{-17}

$\mathcal{O}(10^{-17})$

x10,000 improvement

COMET Collaboration



4



PI: Y. Kuno

The COMET Collaboration

R. Abramishvili¹¹, G. Adamov¹¹, R. Akhmetshin^{6,31}, V. Anishchik⁴, M. Aoki³², Y. Arimoto¹⁸, I. Bagaturia¹¹, Y. Ban³, A. Bondar^{6,31}, Y. Calas⁷, S. Canfer³³, Y. Cardenas⁷, S. Chen²⁸, Y. E. Cheung²⁸, B. Chiladze³⁵, D. Clarke³³, M. Danilov^{15,26}, P. D. Dauncey¹⁴, J. David²³, W. Da Silva²³, C. Densham³³, G. Devidze³⁵, P. Dornan¹⁴, A. Drutskoy^{15,26}, V. Duginov¹⁶, L. Epshteyn^{6,30}, P. Evtoukhovich¹⁶, G. Fedotov^{6,31}, M. Finger⁸, M. Finger Jr⁸, Y. Fujii¹⁸, Y. Fukao¹⁸, J-F. Genat²³, E. Gillies¹⁴, D. Grigoriev^{6,30,31}, K. Gritsay¹⁶, E. Hamada¹⁸, R. Han¹, K. Hasegawa¹⁸, I. H. Hasim³², O. Hayashi³², Z. A. Ibrahim²⁴, Y. Igarashi¹⁸, F. Ignatov^{6,31}, M. Iio¹⁸, M. Ikeno¹⁸, K. Ishibashi²², S. Ishimoto¹⁸, T. Itahashi³², S. Ito³², T. Iwami³², X. S. Jiang², P. Jonsson¹⁴, V. Kalinnikov¹⁶, F. Kapusta²³, H. Katayama³², K. Kawagoe²², N. Kazak⁵, V. Kazanin^{6,31}, B. Khazin^{6,31}, A. Khvedelidze^{16,11}, T. K. Ki¹⁸, M. Koike³⁹, G. A. Kozlov¹⁶, B. Krikler¹⁴, A. Kulikov¹⁶, E. Kulish¹⁶, Y. Kuno³², Y. Kuriyama²¹, Y. Kurochkin⁵, A. Kurup¹⁴, B. Lagrange^{14,21}, M. Lancaster³⁸, M. J. Lee¹², H. B. Li², W. G. Li², R. P. Litchfield³⁸, T. Loan²⁹, D. Lomidze¹¹, I. Lomidze¹¹, P. Loveridge³³, G. Macharashvili³⁵, Y. Makida¹⁸, Y. Mao³, O. Markin¹⁵, Y. Matsumoto³², T. Mibe¹⁸, S. Mihara¹⁸, F. Mohamad Idris²⁴, K. A. Mohamed Kamal Azmi²⁴, A. Moiseenko¹⁶, Y. Mori²¹, M. Moritsu³², E. Motuk³⁸, Y. Nakai²², T. Nakamoto¹⁸, Y. Nakazawa³², J. Nash¹⁴, J. -Y. Nief⁷, M. Nioradze³⁵, H. Nishiguchi¹⁸, T. Numao³⁶, J. O'Dell³³, T. Ogitsu¹⁸, K. Oishi²², K. Okamoto³², C. Omori¹⁸, T. Ota³⁴, J. Pasternak¹⁴, C. Plostinar³³, V. Ponariadov⁴⁵, A. Popov^{6,31}, V. Rusinov^{15,26}, A. Ryzhenkov^{6,31}, B. Sabirov¹⁶, N. Saito¹⁸, H. Sakamoto³², P. Sarin¹³, K. Sasaki¹⁸, A. Sato³², J. Sato³⁴, Y. K. Semertzidis^{12,17}, D. Shemyakin^{6,31}, N. Shigyo²², D. Shoukavy⁵, M. Slunicka⁸, A. Straessner³⁷, D. Stöckinger³⁷, M. Sugano¹⁸, Y. Takubo¹⁸, M. Tanaka¹⁸, S. Tanaka²², C. V. Tao²⁹, E. Tarkovsky^{15,26}, Y. Tevzadze³⁵, T. Thanh²⁹, N. D. Thong³², J. Tojo²², M. Tomasek¹⁰, M. Tomizawa¹⁸, N. H. Tran³², H. Trang²⁹, I. Trekov³⁵, N. M. Truong³², Z. Tsamalaidze^{16,11}, N. Tsverava^{16,35}, T. Uchida¹⁸, Y. Uchida¹⁴, K. Ueno¹⁸, E. Velicheva¹⁶, A. Volkov¹⁶, V. Vrba¹⁰, W. A. T. Wan Abdullah²⁴, M. Warren³⁸, M. Wing³⁸, T. S. Wong³², C. Wu^{2,28}, H. Yamaguchi²², A. Yamamoto¹⁸, Y. Yang²², W. Yao², Y. Yao², H. Yoshida³², M. Yoshida¹⁸, Y. Yoshii¹⁸, T. Yoshioka²², Y. Yuan², Y. Yudin^{6,31}, J. Zhang², Y. Zhang², K. Zuber³⁷

COMET Collaboration



PI: Y. Kuno

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about 200 collaborators
40 institutes, 17 countries

J-PARC@Tokai

COMET
Exp. Area

Hadron Experimental Hall



COMET White Paper

The 2020 Update of
the European Strategy
for Particle Physics

16 December 2018

Further
improvements by
one order of
magnitude

$\mathcal{O}(10^{-18})$

COMET

J.-C. Angélique, C. Cârloganu, W. da Silva, A. Drutskoy, M. Finger,
D. N. Grigoriev, T. Kachelhoffer, F. Kapusta, Y. Kuno¹, P. Lebrun,
R. P. Litchfield, D. Lomidze, D. Shoukavy, A. M. Teixeira, I. Tevzadze,
Z. B. Tsamalaidze, Y. Uchida, V. Vrba, K. Zuber

A submission to the 2020 update of the European Strategy for Particle
Physics on behalf of the COMET collaboration.

Abstract

The search for charged lepton flavour violation (CLFV) has enormous discovery potential in probing new physics Beyond the Standard Model (BSM). The observation of a CLFV transition would be an undeniable sign of the presence of BSM physics which goes beyond non-zero masses for neutrinos. Furthermore, CLFV measurements can provide a way to distinguish between different BSM models, which may not be possible through other means. So far muonic CLFV processes have the best experimental sensitivity because of the huge number of muons which can be produced at several facilities world-wide, and in the near future, new muon beam-lines will be built, leading to increases in beam intensity by several orders of magnitude. Among the muonic CLFV processes, $\mu \rightarrow e$ conversion is one of the most important processes, having several advantages compared to other such processes.

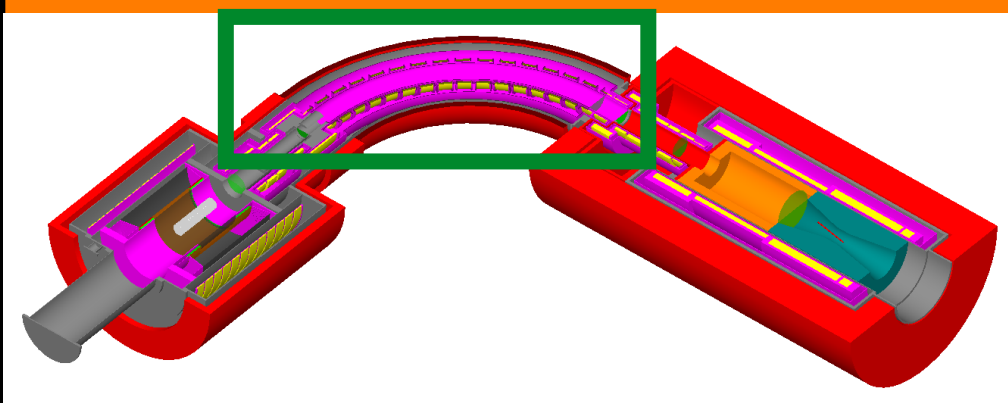
We describe the COMET experiment, which is searching for $\mu \rightarrow e$ conversion in a muonic atom at the J-PARC proton accelerator laboratory in Japan. The COMET experiment has taken a staged approach; the first stage, COMET Phase-I, is currently under construction at J-PARC, and is aiming at a factor 100 improvement over the current limit. The second stage, COMET Phase-II is seeking another 100 improvement (a total of 10,000), allowing a single event sensitivity (SES) of 2.6×10^{-17} with 2×10^7 seconds of data-taking. Further improvements by one order of magnitude, which arise from refinements to the experimental design and operation, are being considered whilst staying within the originally-assumed beam power and beam time. Such a sensitivity could be translated into probing many new physics constructions up to $\mathcal{O}(10^4)$ TeV energy scales, which would go far beyond the level that can be reached directly by collider experiments. The search for CLFV $\mu \rightarrow e$ conversion is thus highly complementary to BSM searches at the LHC.

¹contact person: kuno@phys.sci.osaka-u.ac.jp.

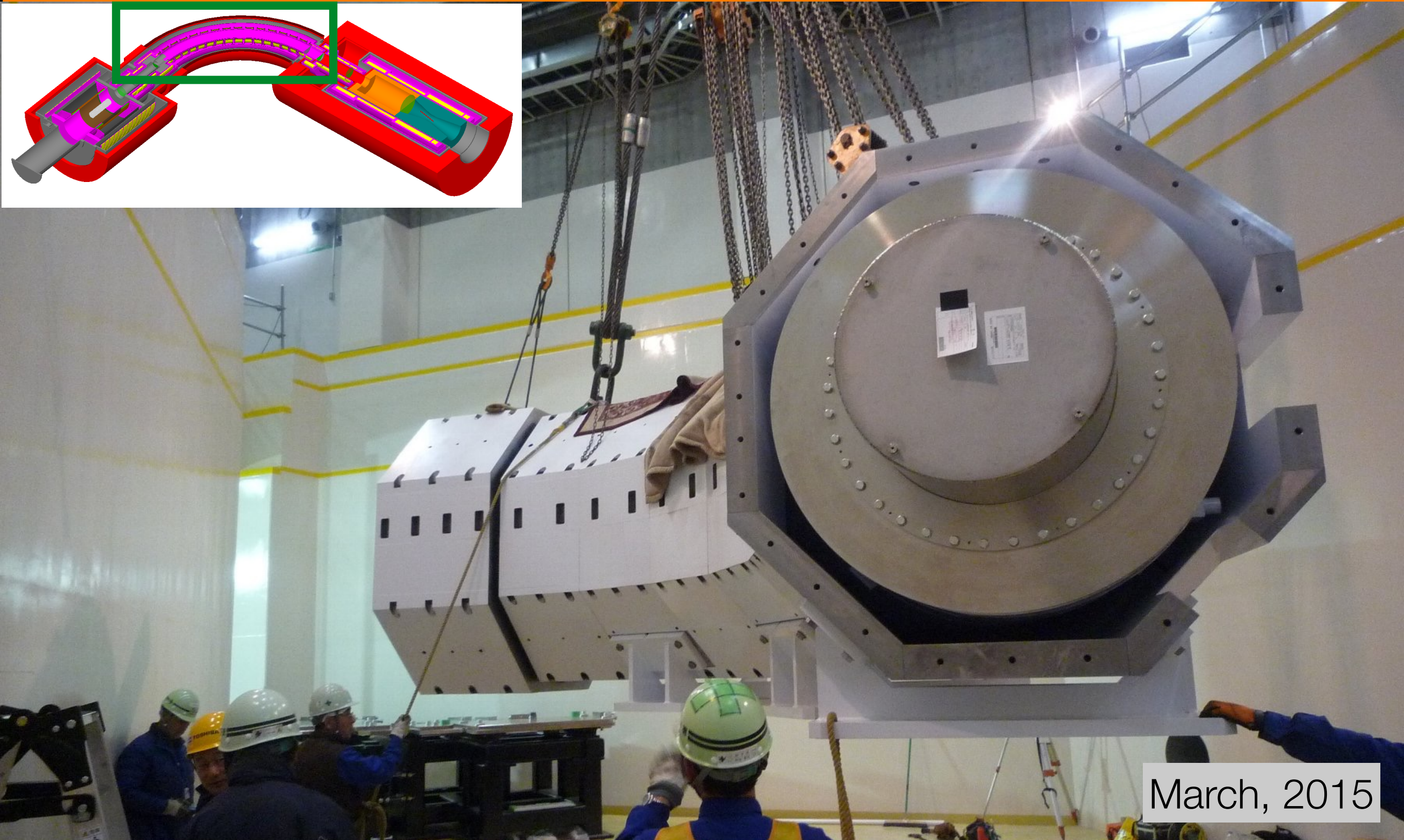
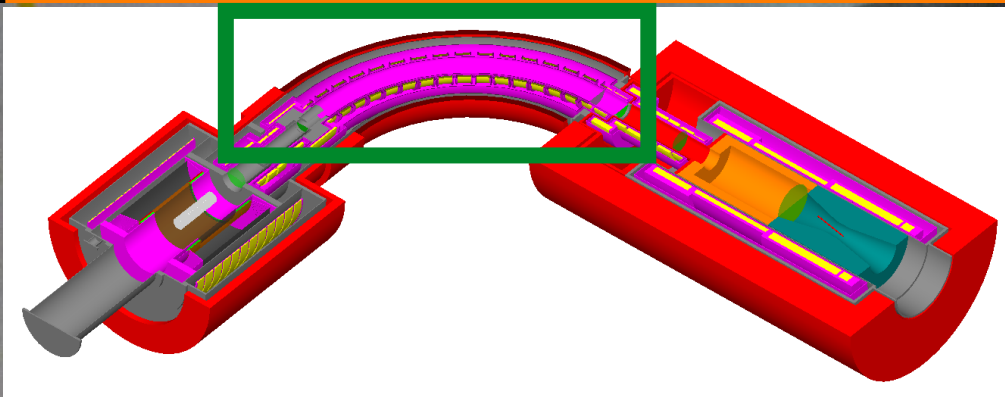
Curved Solenoids for Muon Transport Completed and Delivered!



Curved Solenoids for Muon Transport Completed and Delivered!

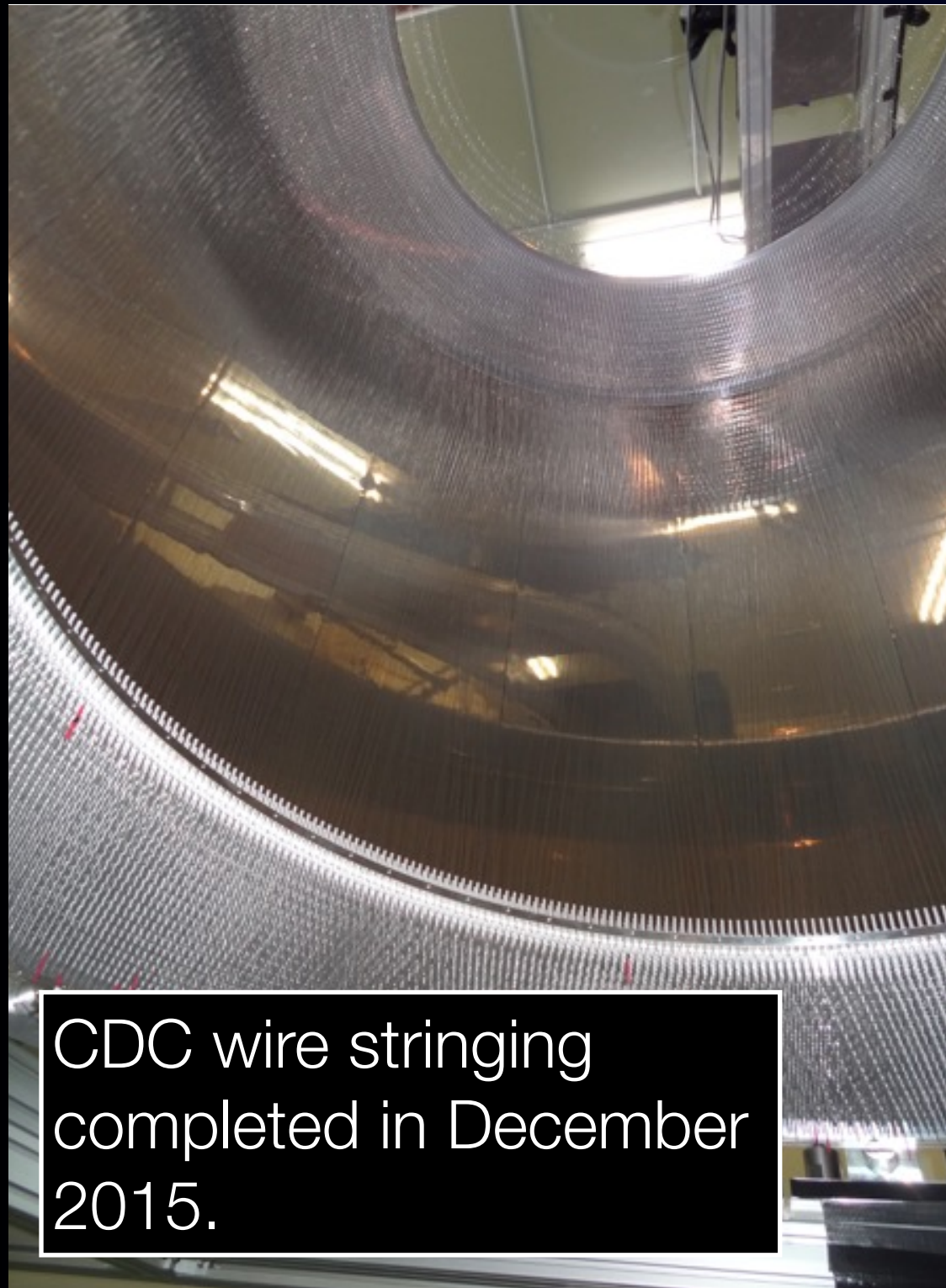


Curved Solenoids for Muon Transport Completed and Delivered!



March, 2015

CDC Construction completed!

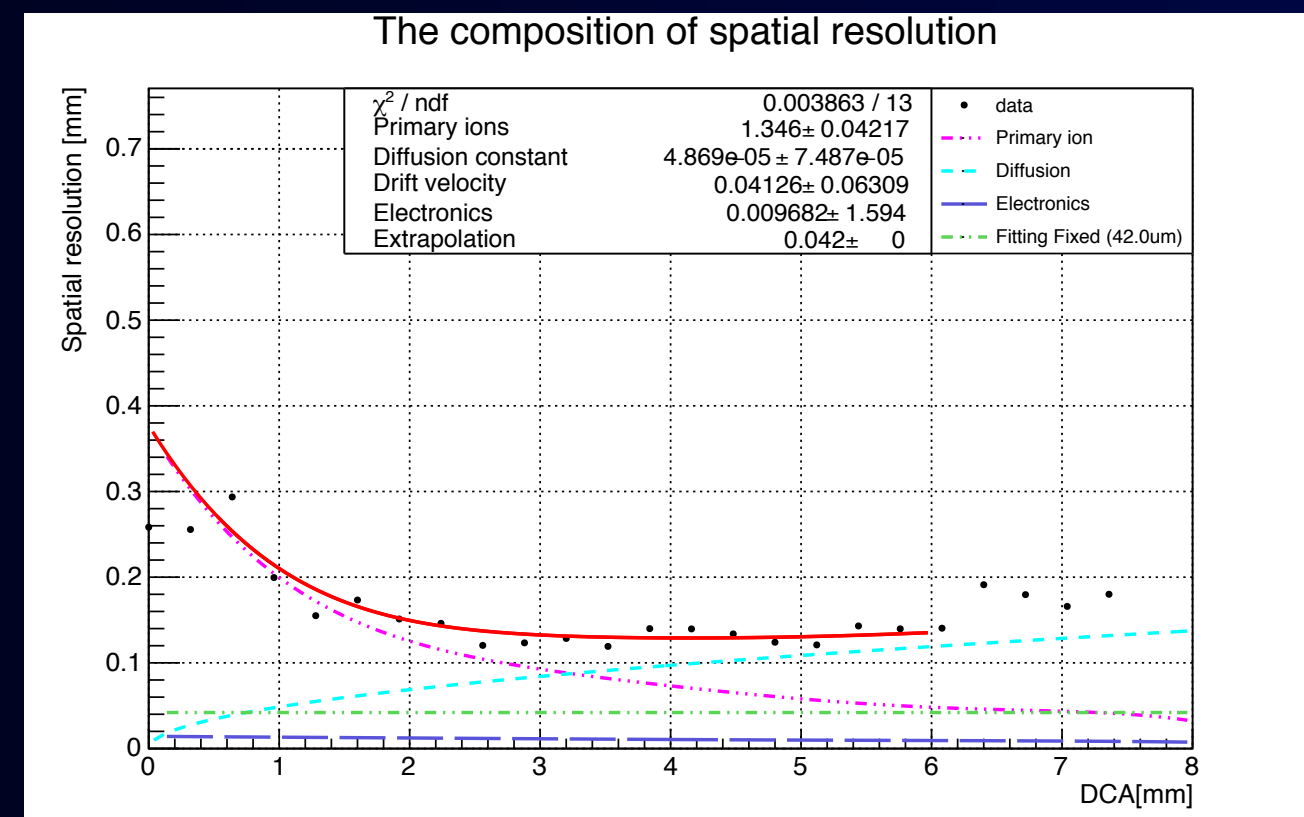
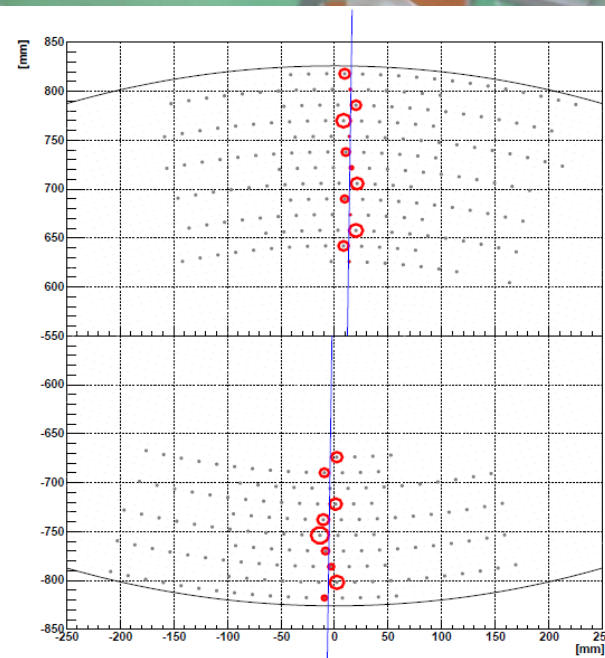
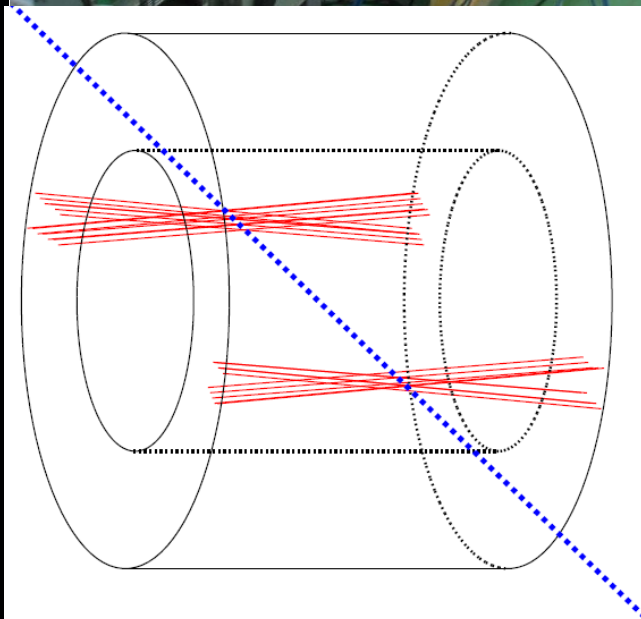
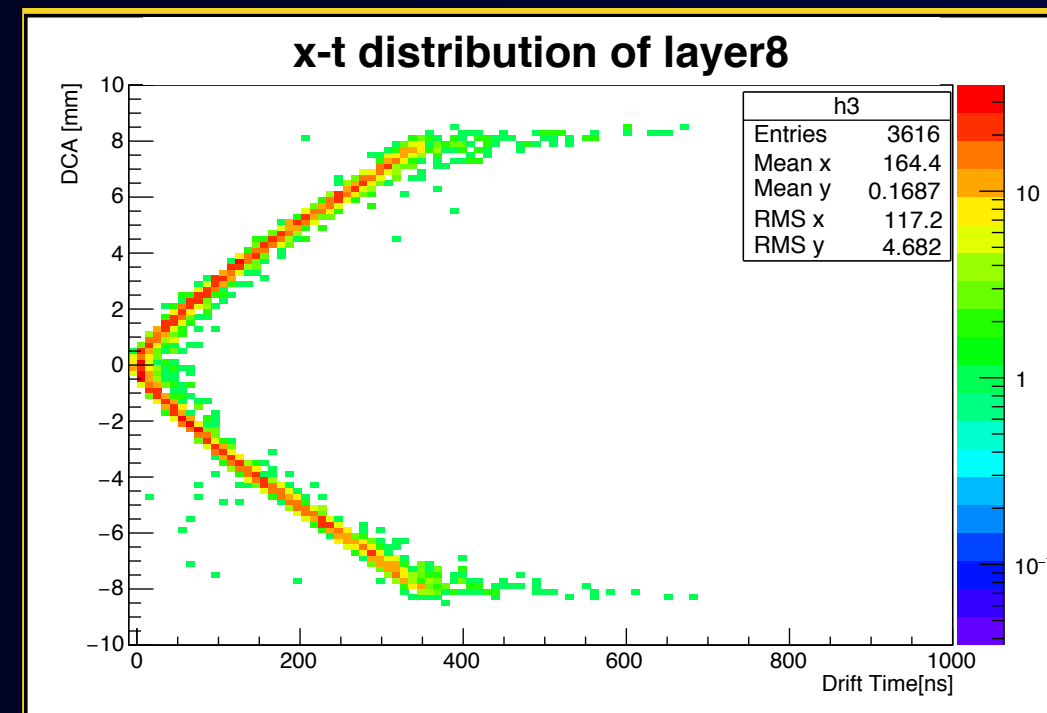
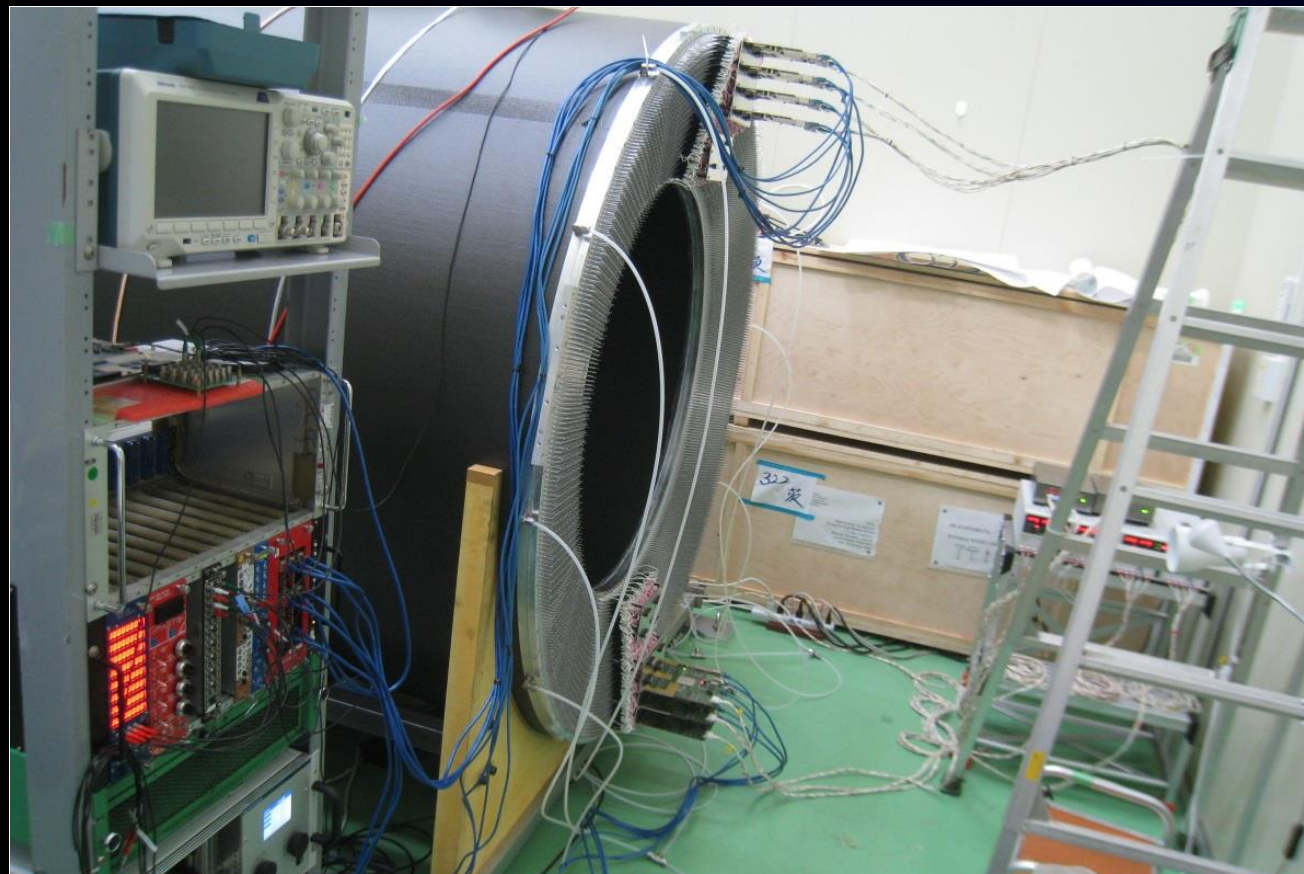


CDC wire stringing completed in December 2015.



CDC readout electronics completed in August 2015.

CDC Cosmic Ray Test

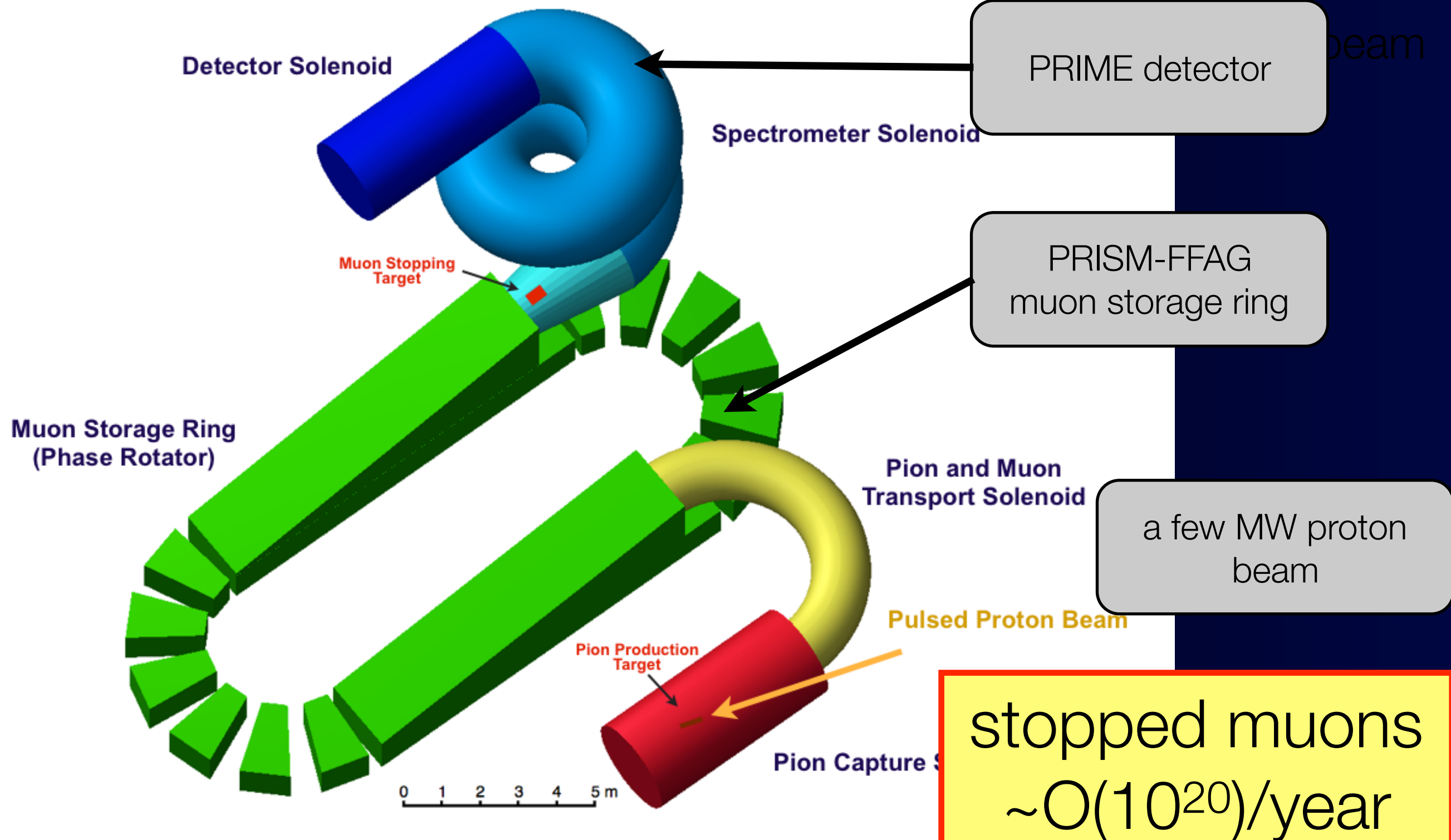


$\mu - e$ conversion

PRISM-PRIME@J-PARC

PRISM (=Phase Rotated Intense Slow Muon source)

PRISM/PRIME



R&D on the PRISM-FFAG Muon Ring at Osaka University



PRISM-FFAG (6 sectors) in RCNP, Osaka



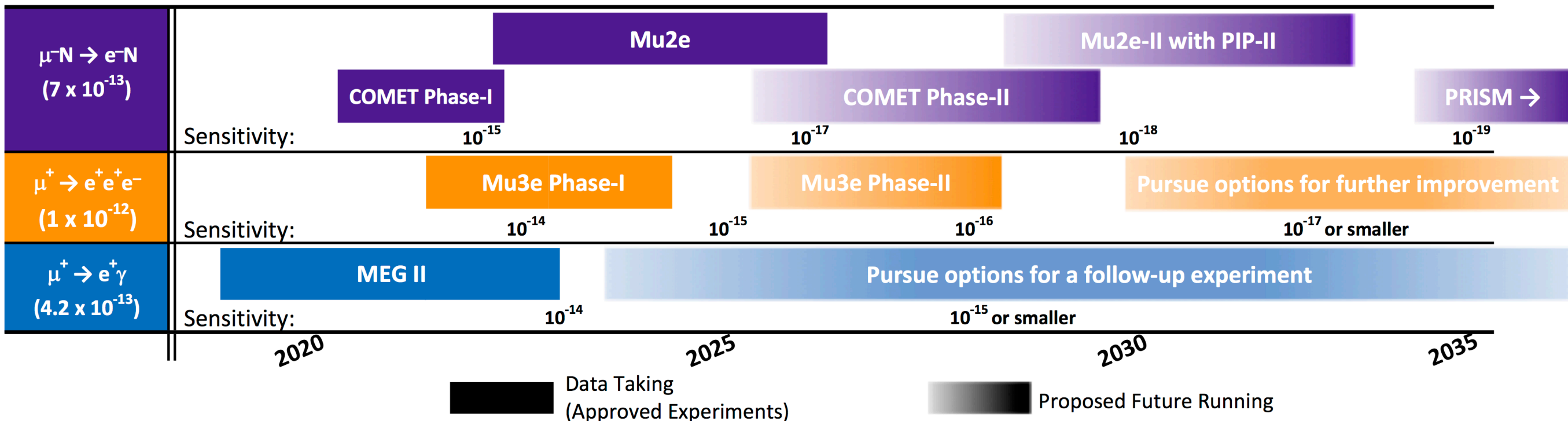
demonstration of phase rotation has been made.

Global Timeline of Muon CLFV



from the muon CLFV white paper for the 2020
update of European Strategy of Particle Physics

Searches for Charged-Lepton Flavor Violation in Experiments using Intense Muon Beams



COMET Phase-I

$\mathcal{O}(10^{-15})$

COMET Phase-II

$\mathcal{O}(10^{-18})$

PRISM

$\mathcal{O}(10^{-19})$

Summary



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- The muon sources will be improved from 10^{18} /year (1000 more)
- Major muon CLFV processes are $\mu \rightarrow e\gamma$, $\mu \rightarrow eee$, $\mu \rightarrow e$ conversion, whilst there are other interesting muon CLFV processes.
- COMET @ J-PARC is seeking for $\mu \rightarrow e$ conversion with sensitivity of $O(10^{-18})$ with staged approach, 100,000 improvement.
- COMET Phase-I with 100 improvement is under preparation, starting soon.

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my dog, IKU



Thank you for
your attention!